

A New Code for Spectrometric Analysis for Environmental Radiological Surveillance on Monitors Focused on Gamma Radioactivity on Aerosols

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Abstract—This paper presents a new code for the analysis of gamma spectra generated by an equipment for continuous measurement of gamma radioactivity in aerosols with paper filter. It is called pGamma and has been developed by the Nuclear Engineering Research Group at the Technical University of Catalonia - Barcelona Tech and by Raditel Serveis i Subministraments Tecnològics, Ltd. The code has been developed to identify the gamma emitters and to determine their activity concentration. It generates alarms depending on the activity of the emitters and elaborates reports. Therefore it includes a library with NORM and artificial emitters of interest. The code is being adapted to the monitors of the Environmental Radiological Surveillance Network of the local Catalan Government in Spain (Generalitat de Catalunya) and is used at three stations of the Network.

Index Terms—Environmental radiation, gamma activity, gamma spectrometry, isotope identification.

I. INTRODUCTION

THE use of monitors with spectrometric capability in environmental radiological surveillance networks provides additional information that complements the ambient dose equivalent $H^*(10)$. Based on this premise, our group (Nuclear Engineering Research Group) at the Technical University of Catalonia - Barcelona Tech and Raditel Serveis i Subministraments Tecnològics, Ltd. have developed an equipment for continuous measurement and identification of gamma radioactivity in aerosols with paper filter. Gamma spectrum analysis allows the identification and determination of activity concentration of radiation sources. Excellent commercial software is available for this purpose (among others, GammaVision from ORTEC and Genie2000 from

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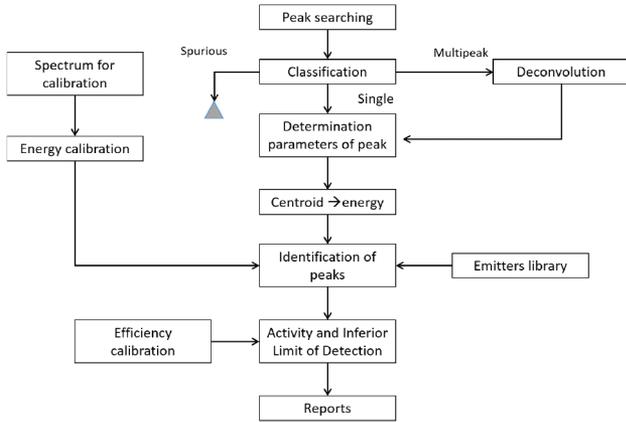
CANBERRA). Nevertheless, it is generic software not intended for automatic analysis since each individual analysis must be conducted by a specialised technician. For this reason, our research group has developed a spectrometric analysis code specifically designed for the above equipment. This code is being used at stations with aerosol monitors on the Environmental Radiological Surveillance Network of the Generalitat de Catalunya (local Catalan Government), Spain¹: Three monitors are currently fully operative: two in the surveillance station of the Ascó and Vandellòs Nuclear Power Plants (both in the province of Tarragona, Spain), and one in the Engineering School of Barcelona, Spain. Two new monitors will shortly be deployed in Roses (province of Girona) and in Puigcerdà (province of Barcelona). The code has been adapted for the analysis of gamma spectra generated by other monitors with spectrometric capability of the Network, i.e. river and direct air monitors. We are starting the tuning of these versions of the code on the monitors for river and direct air.

The code, called pGamma, discriminates spectrum information, identifies emitters appearing on the spectrum, determines activity concentration in Bq/m^3 and generates alarms according to its calculated values. It has a normal operation mode and special operation mode for spectra obtained during energy calibration of the equipment. Additionally, another mode called investigation, is set when the variation in the total number of counts exceeds a certain level, or when a full energy peak is not identified.

The code input is an ASCII text file with the number of counts per channel of the spectrum and a header with information about measurement conditions: date and time of counting initiation and termination, live time, air flow, operation mode, the parameters of the energy calibration curve, etc. Fig. 1 is a general operation diagram of pGamma in normal operation mode.

The paper is organized as follows: Section II describes the treatment of the peaks in the spectrum. In Section III, efficiency calibration is presented. The determination procedure of activity concentration is shown in Section IV. Section V describes the identification of ¹³¹I. In Section VI, the action levels are

¹This network is integrated in the Spanish Nuclear Security Council network.



F1:1 Fig. 1. General diagram of pGamma.

77 presented. An Ascó station spectrum is analysed in Section VII.
78 Finally, conclusions are drawn in Section VIII.

79 II. PEAK TREATMENT

80 The first step in the analysis of the spectrum is the treatment
81 of the peaks: the searching of possible peaks on the spectrum
82 and its identification.

83 A. Peak Search

84 Peak search and analysis is performed after the input spec-
85 trum is smoothed using the Savitzky-Golay method [1]. The
86 smoothing algorithm allows the first, second and third deriva-
87 tives of the spectrum to be obtained. The first derivative is used
88 to determine the local maximums, which are subjected to a sta-
89 tistical test to discriminate real center of peaks (centroids) from
90 spurious transitions. The nearest relative maximums of the first
91 derivative at both sides of the centroid, indicate the channel of
92 the peak boundaries. Using the second and the third derivate,
93 single peaks are discriminated from multiplets. With the cen-
94 troid and the boundaries of the peaks, the net area and the
95 background are determined. The last step is the fitting of each
96 peaks to a Gaussian curve. It allows the determination of the full
97 width at half maximum (FWHM) and perform the chi-square
98 test for further discrimination. With this information, pGamma
99 generates a first list of possible peaks with their characteris-
100 tics. The calculated parameters are centroid, lower and upper
101 boundaries of the possible peak, full width at half maximum,
102 gross area and net area.

103 B. Peak Identification

104 1) *Energy Calibration:* The use of energy calibration
105 makes it possible to state the centroid, FWHM and limits
106 of peaks in energy units instead of number of channels. The
107 relationship between energy and adopted number of channel is:

$$E = a + bC + cC^2 \quad (1)$$

108 Where C is the channel and a , b and c are the parameters of the
109 calibration curve. Energy calibration must be performed prior
110 to installation of equipment in a station.

2) *Peak Identification:* Peaks can be identified using the
111 library of gamma emitters of interest included in the code.
112 This library has been created by us using information from [2].
113 For aerosol equipment, these emitters are the NORM radionu-
114 clides (^{238}U - ^{222}Rn series, ^{232}Th - ^{220}Rn series, ^{40}K and ^7Be)
115 and artificial emitters from nuclear power plant discharges
116 (e.g. radio-iodines, ^{137}Cs , ^{134}Cs and ^{132}Te) or industrial appli-
117 cations. Information about emitters includes energy and the
118 emission probability of emitted gamma rays, empirical thresh-
119 olds for NORM emitters to decide if pGamma must switch to
120 investigation mode and derived concentration limits in the air.
121

The centroid energy of each peak is compared with the
122 energy of gamma emitters. For similar values, the emitters are
123 considered candidates for that peak and a list of candidates
124 is generated. To identify a set of peaks with a radionuclide
125 with multiple gamma emissions, the ratio between intensity
126 of detected gammas and total intensity of the gammas of the
127 emitter in the library is used. The relationship between the
128 members of ^{238}U - ^{222}Rn series and ^{232}Th - ^{220}Rn series is con-
129 sidered too. pGamma cannot determine activity concentration
130 of unidentified peaks.
131

132 III. DETECTION EFFICIENCY CALIBRATION

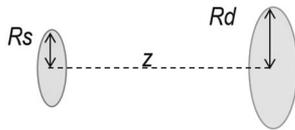
In order to determine activity concentration, the code needs
133 information about the detection efficiency. Since the factors
134 integrating the detection efficiency are very difficult to deter-
135 mine due to a complex geometry and a wide range of possible
136 energies, efficiency calibration must be performed. First, the
137 same geometry of the paper filter (source), detector and detec-
138 tion volume of the equipment during normal operation must
139 be reproduced. This implies preparing a calibration source
140 by shaping it into a disc with the same diameter of aerosols
141 retained on the filter, and the same self-absorption (negligi-
142 ble) and gamma emitters of several energies in the range from
143 100 keV to 2000 keV. In consequence, the analytical determi-
144 nation of efficiency is cumbersome. The solution consists in
145 using a calibration source with similar characteristics to the
146 source to be measured and determining efficiency from experi-
147 mental values. To ensure unbiased calibration, the calibration
148 source should be identical to the radioactive samples in all
149 aspects affecting detection [3]. If no calibration source is avail-
150 able or more points are required for adjusting the efficiency
151 curve, Monte Carlo n-Particle (MCNP5, version 2.6.0) provide
152 a further possibility for calculating the detector efficiency.
153

We used simulated ^{131}I and commercial ^{60}Co sources
154 (Table I) with similar geometry and materials to those of the
155 impregnated filter of the monitor. The simulated ^{131}I source
156 is composed of two radioisotopes: ^{133}Ba and ^{137}Cs . The com-
157 bination of both emitters generates a gamma energy spectrum
158 whose full energy peaks have centroids and a relationship of
159 areas very similar to the ^{131}I spectrum. The simulated ^{131}I
160 source has the same area as the impregnated filter. The emitter
161 material is deposited on a polymeric membrane with a stainless
162 steel backing of 0.762 mm thickness and an 0.9 mg/cm^2 alu-
163 minized mylar window. Gamma absorption of the window and
164 the backscattering on the back of the membrane are negligible.
165 Therefore, it can be considered that the material and geometry
166

T1:1
T1:2
T1:3

TABLE I
CHARACTERISTICS OF THE CALIBRATION SOURCES (REFERENCE FOR
ACTIVITY: 01-JANUARY-2014)

(a)	
Nuclides	^{137}Cs , ^{133}Ba
^{137}Cs Activity	0.4096 kBq
^{133}Ba Activity	4.222 kBq
^{131}I Activity	3.684 kBq
Manufacturer	Eckert & Ziegeler
Model	EAB-131(s)-PL
Source No.	K9-928
(b)	
Activity	3.863 kBq
Manufacturer	Eckert & Ziegeler
Model	EAB-060-PL
Source No.	K9-925



F2:1 Fig. 2. Geometry for equations (2) and (4).

167 of the source have the same behaviour as the impregnated paper
168 filter. To simulate the detector response on the MCNP5 sim-
169 ulation we had used the FWHM obtained from the peaks on
170 the spectra for the calibration sources (^{131}I and ^{60}Co) with the
171 experimental setup. The function used is:

$$FWHM = 0.008 + 0.013 \cdot \sqrt{E + 0.593 \cdot E^2} \quad (2)$$

172 where E is the energy in MeV and the FWHM obtained is also
173 in MeV.

174 Experimental efficiencies were obtained by analyzing the
175 experimental spectra of ^{131}I and ^{60}Co by pGamma. Before the
176 calculation of the efficiency, the net counts of the experimental
177 peaks had to be compensated by a correction factor. The source
178 capsule elevated the membrane where the radioactive material
179 is deposited such that the membrane is slightly higher than the
180 paper filter (about 0.283 cm). Assuming that both sources, the
181 impregnated paper filter in normal operation and the calibration
182 source, are discs and the detector window is a disc too (geome-
183 try on Fig. 2), we could determine the geometric factor in both
184 cases using the following expansion in power series:

$$gf_i = \frac{1}{2} \left(1 - \frac{1}{R} \right) - \frac{3 \cdot \omega^2 \psi^2}{16 \cdot R^5} \cdot \left[1 - \frac{5 \psi^2}{6 \cdot R^4} \cdot \left(1 - \frac{3 \cdot \omega^2}{4} \right) + \frac{35 \cdot \psi^4}{48 \cdot R^8} \left(1 - \frac{5 \cdot \omega^2}{2} + \frac{5 \cdot \omega^4}{8} \right) + \dots \right] \quad (3)$$

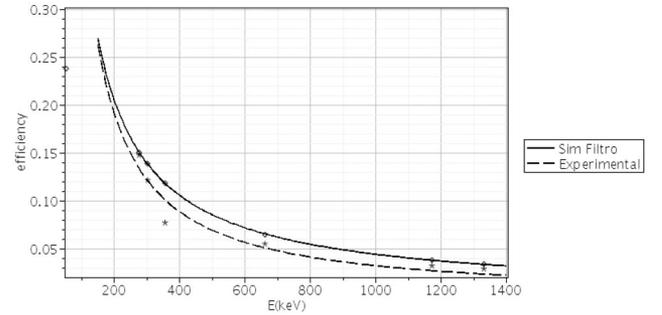
185 the factors R , ψ and ω are determined as:

$$\psi = \frac{R_S}{z}; \omega = \frac{R_d}{z}; R = \sqrt{1 + \omega^2} \quad (4)$$

186 where R_S is the source radius, R_d is the detector radius, z is the
187 distance between the source and the detector. Equation (4) can

TABLE II
EXPERIMENTAL DETECTION EFFICIENCY

Peak	E_i (keV)	N_i (cps)	S_i (ph/s)	ε_d
^{131}I sim	53.160	57.990	88.510	0.70319
^{131}I sim	276.390	40.801	296.150	0.14786
^{131}I sim	303.000	86.024	758.550	0.12170
^{131}I sim	356.010	184.860	2566.500	0.017730
^{131}I sim	661.500	19.830	385.100	0.05527
^{60}Co	1173.000	108.028	3573.290	0.03244
^{60}Co	1332.000	97.690	3579.020	0.02929



F3:1 Fig. 3. Curve of efficiency.

be applied either for the calibration source or the impregnated
188 filter. Then the correction factor cf is:
189

$$cf = \frac{gf_{calibration\ source}}{gf_{filter}} \quad (5)$$

where gf is the geometric factor. With this method, the cor-
190 rection factor is 0.93680. By using seven MCNP5 code sim-
191 ulations, one for each energy, the mean correction factor is
192 0.93184. Efficiency ε_d for energy E_i is determined as:
193

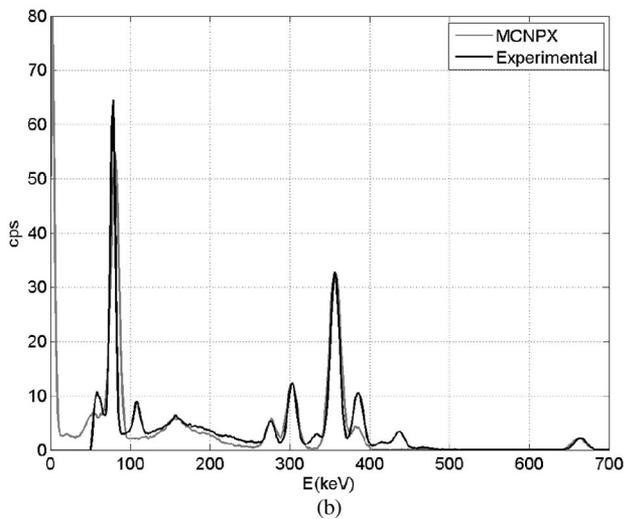
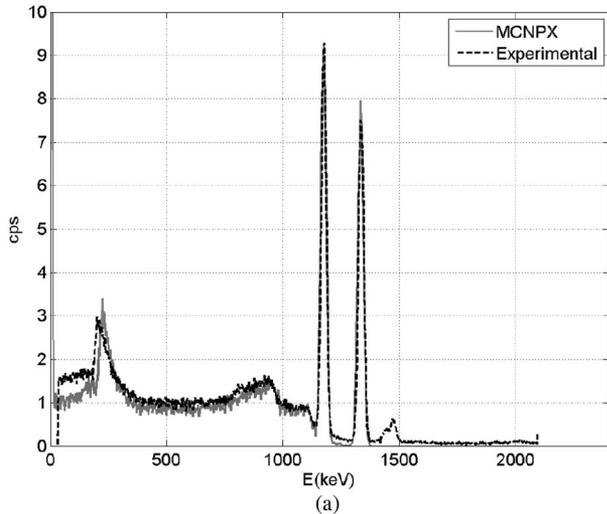
$$\varepsilon_d = \frac{1}{cf} \cdot \frac{N_i}{S_i} \quad (6)$$

where N_i represents the net count rate of the full energy peak
194 corresponding to the source gamma rays with energy E_i and S_i
195 is the intensity of photons with energy E_i . Intensity is deter-
196 mined for the two components of the simulated ^{131}I : ^{137}Cs and
197 ^{133}Ba , which are treated as two separate sources.
198

$$S_i = \nu_i \cdot A_i(0) e^{-T_t \cdot \lambda_i}; i = ^{137}\text{Cs}, ^{133}\text{Ba} \quad (7)$$

where ν_i is the emission probability of photons with energy E_i ,
200 $A_i(0)$ is the activity on the calibration date of the source, T_t is
201 the time between the calibration date of the source and the day
202 of the experiment and λ_i is the disintegration constant. With
203 equations (6) and (7) we find the values of ε_d on Table II. A
204 second efficiency curve, obtained from MCNP5, is compared
205 with that obtained from experimental values, Fig. 3. As can
206 be seen, the two curves are very close, with the exception of
207 experimental efficiency for 380 keV. By fitting the experimen-
208 tal points without considering the 380 keV experimental value,
209 efficiency is:
210

$$\varepsilon_d = \begin{cases} 0.43792 & 100 \text{ keV} \leq E \leq 200 \text{ keV} \\ 60.71703 \cdot E^{-1.07096} & E > 200 \text{ keV} \end{cases} \quad (8)$$



F4:1 Fig. 4. Comparison between simulated and experimental spectra. Experimental
 F4:2 setup: $\text{LaBr}_3 2'' \times 2''$ scintillator and the DigiBASE compact instrumentation
 F4:3 from ORTEC. The source is inside the volume of detection equipment for con-
 F4:4 tinuous measurement and identification of gamma radioactivity in aerosols with
 F4:5 paper filter. (a) Co-60 (b) Simulated I-131.

211 Since Monte Carlo methods were used to complement the
 212 experimental values to obtain the efficiency curve, the equip-
 213 ment simulation model must be benchmarked. In fact, without
 214 the geometry correction, Fig. 3 itself could be a validation of
 215 the Monte Carlo simulations. It is interesting to analyze and
 216 compare the experimental and Monte Carlo simulated spec-
 217 tra of the equipment with the calibration source instead of the
 218 paper filter. As the sources were very close to the front win-
 219 dows of the detector, some true coincidence summing peaks
 220 appeared on the spectrum. These cannot be reproduced with
 221 a single MCNP5 simulation, but as we want to compare the
 222 procedure used to obtain the efficiency curve and these sum
 223 peaks are not required for the efficiency calculations, we will
 224 not simulate them. Fig. 4 compares simulated and experimen-
 225 tal ^{60}Co and simulated source of ^{131}I spectra. The simulation
 226 reproduces perfectly the processes that not depend on the
 227 electronic instrumentation modules. As can be seen, backscat-
 228 tering peaks, Compton continuum, Compton valley and two

TABLE III
 COMPARISON OF NET COUNTS OF PEAKS

Net Counts ^{60}Co (cps)				
E (keV)	Experim.	MCNP		
1173	107.95	122.13		
1332	97.62	112.34		

Net Counts ^{131}I (cps)		Gross counts ^{131}I (cps)		
E (keV)	Experim.	MCNPX	Experim.	MCNP
60	24.11	8.60	57.99	65.19
80	189.18	303.36	260.91	380.45
276	21.02	17.61	40.80	44.48
303	63.31	73.27	86.02	104.65
356	184.86	249.07	231.13	291.88
661.5	19.83	22.51	22.32	22.79

229 full energy peaks appear. The differences between both spectra
 230 are due to the fact that background radiation is not simulated.
 231 Intrinsic radiation from the $\text{LaBr}_3(\text{Ce})$ scintillator, i.e. ^{138}La
 232 and ^{138}La plus X-rays peaks at 1468 keV, appears in the experi-
 233 mental spectrum. This component was not simulated, but is not
 234 relevant for efficiency calibration. Moreover, at low energies
 235 the simulated spectrum has fewer counts than the experimental
 236 one. This is because the Monte Carlo model does not simulate
 237 the background radiation components of cosmic rays or earth's
 238 NORMs, resulting in fewer X-rays on the lead shielding (peak
 239 near 80 keV in the spectrum). In our model, only X-rays from
 240 absorption of the calibration source radiation are generated.
 241 Again, this is not important for efficiency calculation.

242 Since the simulated ^{131}I has many peaks at low energies, the
 243 effect of sum peaks is more noticeable, at least in our energy
 244 range (0-2048 keV). Monte Carlo does not generate sum peaks,
 245 as can be seen in Fig. 4. For full energy peaks with little or no
 246 influence of sum peaks, simulation results are good. Table III
 247 compares the net areas of full energy peaks of simulated and
 248 experimental spectra. Peaks influenced by a sum peak are not
 249 used in efficiency calibration.

IV. DETERMINATION OF THE ACTIVITY CONCENTRATION 250

251 After identification of gamma emitters in the spectrum, the
 252 spectrometric analysis system determines their specific activity
 253 in the air \tilde{A}_a , expressed in Bq/m^3 , according to the net area of
 254 the most efficient emitter of the isotope. However, the obtained
 255 spectrum corresponds to the number of counts caused by the
 256 radionuclide concentration on the filter, which is different from
 257 the radionuclide concentration in the air. Because the activity on
 258 the filter A_f is proportional to the number of captured atoms,
 259 a balance of activity can be performed in the same way as for
 260 concentration. The radionuclide concentration of an analysed
 261 emitter on the filter at a certain time depends on three factors:
 262 1) the number of captured atoms of this emitter; 2) decays of
 263 the parent nuclei of this emitter, and 3) decays of this emitter.

264 Activity on the filter A_f is proportional to activity in the air
 265 \tilde{A}_a , air flow Q , and retention efficiency of the filter ε_f [4], [5].
 266 Nevertheless, as activity on the filter implies a balance of matter
 267 on the filter, concentration of each radionuclide depends on its

268 position in the decay series. In the case of aerosol equipment,
269 the radionuclides are belonging to the ^{238}U - ^{222}Rn series, the
270 ^{232}Th - ^{220}Rn (chiefly ^{212}Pb , ^{212}Bi and ^{208}Tl) series or to none
271 of the series. A slightly different statement is required for each
272 case.

273 A. Radionuclides not Belonging to a Series

274 Most artificial emitters fall in this category. The balance of
275 matter in this case is simply:

$$\frac{dA_f}{dt} = Q\varepsilon_f\tilde{A}_a(0) - \lambda_i A_f \quad (9)$$

276 Equation (9) can be solved by integration and as result the filter
277 activity of the emitter analysed is:

$$A_f(t) = \frac{Q\varepsilon_f\tilde{A}_a}{\lambda} \int_0^T 1 - e^{-\lambda t} dt \quad (10)$$

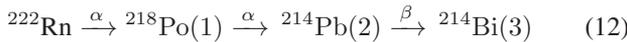
278 Then, from Equation (10), the specific activity \tilde{A}_a in the air for
279 emitter i (expressed Bq/m³) is:

$$\tilde{A}_a = \frac{\lambda_i N_i T}{Q\varepsilon_f\varepsilon_d\nu_i \int_0^T (1 - e^{-\lambda_i t}) dt} \quad (11)$$

280 Where λ_i is the decay constant for radionuclide i , N_i the net
281 area (counts per second) of the full energy peak of the most
282 probable gamma of radionuclide i , T the detection time (live
283 time fo the Multiple Channel Analyzer), Q the average air-
284 flow, ε_f the filter efficiency, ε_d the detection efficiency for the
285 most probable gamma of radionuclide i , and ν_i the emission
286 probability of the most probable gamma of radionuclide i .

287 B. ^{238}U - ^{222}Rn Series

288 The gamma emitters in the air of interest belonging to the
289 decay series of ^{238}U are the descendents of ^{222}Rn :



290 Numbers in parenthesis in the chain (12) represent the nomen-
291 clature used for the radionuclides in the following equations.
292 These radionuclides are present on all countings because they
293 come from the earth and surrounding buildings. Among them
294 we can find some gamma emitters of certain importance, mainly
295 ^{214}Pb and ^{214}Bi . The equations of balance of matter for chain
296 (12) are:

$$\begin{aligned} \frac{dA_{f1}}{dt} &= \tilde{A}_{a1}(0)Q\varepsilon_f - \lambda_1 A_{f1} \\ \frac{dA_{f2}}{dt} &= \lambda_1 A_{f1} + \tilde{A}_{a2}(0)Q\varepsilon_f - \lambda_2 A_{f2} \\ \frac{dA_{f3}}{dt} &= \lambda_2 A_{f2} + \tilde{A}_{a3}(0)Q\varepsilon_d - \lambda_3 A_{f3} \end{aligned} \quad (13)$$

297 ^{218}Po reaches its equilibrium at 95% in 13 minutes. By com-
298 paring the partial time of detection of 1 h and the total time of
299 detection of 24 h, the hypothesis that ^{218}Po and ^{214}Pb are in
300 equilibrium in the air can be supported.

$$\tilde{A}_{a1}(0) = \tilde{A}_{a2}(0) \quad (14)$$

By considering the equilibrium condition Eq. (14) and devel- 301
oping the balance equations Eq. (13), the specific activities of 302
 ^{214}Pb and ^{214}Bi can be determined by: 303

$$\tilde{A}_{a2}(0) = \frac{1}{2} \frac{N_2 T \lambda_2}{Q\nu_2\varepsilon_f\varepsilon_{d2} \int_0^T (1 - e^{-\lambda_2 t}) dt} \quad (15)$$

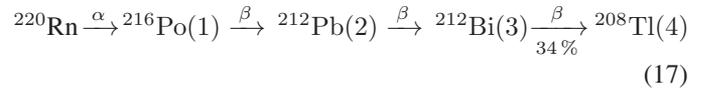
The equation of balance of matter for chain 12 is (for simplicity 304
we omit the explicit time dependences here): 305

$$\begin{aligned} \tilde{A}_{a3}(0) &= \frac{\lambda_3 N_3 T}{Q\nu_3\varepsilon_f\varepsilon_{d3} \int_0^T (1 - e^{-\lambda_3 t}) dt} \\ &- \frac{2\tilde{A}_{a2}(0)}{\int_0^T (1 - e^{-\lambda_3 t}) dt} \left[\int_0^T (1 - e^{-\lambda_3 t}) dt \right. \\ &\left. + \frac{\lambda_3}{\lambda_2 - \lambda_3} \int_0^T (e^{-\lambda_2 t} - e^{-\lambda_3 t}) dt \right] \end{aligned} \quad (16)$$

Emission probability ν_i , the net area of the full energy peak 306
 N_i and detection efficiency ε_i correspond to the most probable 307
gamma of the analysed radionuclide. 308

C. ^{232}Th - ^{220}Rn Series

As in the previous case, only the nuclides at the end of the 310
chain are of interest in this series, i.e. from ^{220}Rn : 311

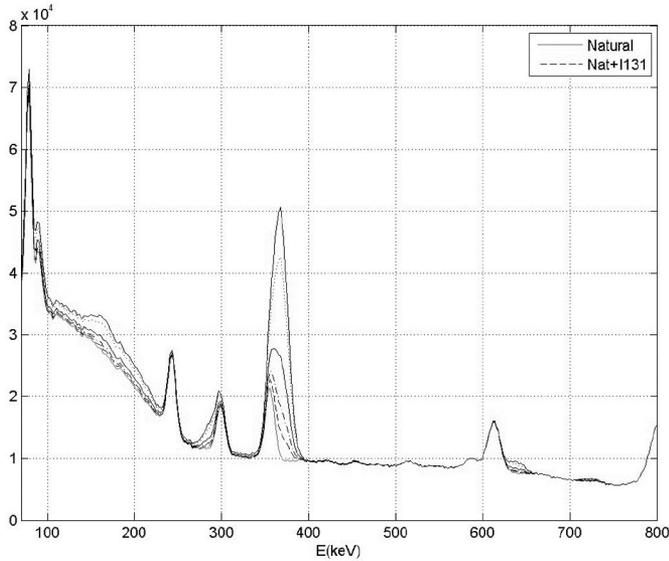


These radionuclides are always present in air aspiration, like 312
in the previous series. Among them, ^{212}Pb , ^{212}Bi and ^{208}Tl 313
are detected by the equipment. ^{212}Pb bifurcates to ^{208}Tl with 314
35.94% and to ^{212}Po with 64.06%, but ^{212}Po is not an impor- 315
tant γ emitter. However, in the location areas of our monitoring 316
stations there are fewer of those than of the ^{238}U - ^{222}Rn series, 317
and in typical spectra generated by the equipment only the pres- 318
ence of ^{212}Pb is detected. The balance of matter is similar to 319
that of the previous series Eq. 13. The half-life of ^{216}Po is 320
0.15 s. Hence, it can be assumed that at the end of detection 321
equilibrium exists between ^{216}Po and ^{212}Bi . 322

V. ANALYSIS OF ^{131}I

A. Identification and Minimum Detectable Activity (MDA)

Full energy peaks of ^{131}I gamma emitters are very close 325
to some full energy peaks of ^{212}Pb , ^{214}Pb and ^{214}Bi . These 326
natural emitters belong to the ^{222}Rn radon and ^{220}Rn thoron 327
series, and are always present in the spectrum. Overlapping 328
of iodine gamma lines with other emitters occurs between 329
the 284.2 keV, 364.4 keV and 636.97 keV peaks of ^{131}I . 330
The 722.89 keV peak does not overlap with those of natu- 331
ral emitters, but it has very low emission probability. Like the 332
80.18 keV of ^{131}I , that does not overlap with other peaks, but is 333
outside the linear operating range of the scintillator. Therefore, 334
to identify ^{131}I we can only use full energy peaks overlapped 335
with natural emitter peaks, which are always present in the 336



F5:1 Fig. 5. Spectra of ^{131}I . Obtained by a combination of experimental back-
 F5:2 ground and simulations with the filter impregnated with ^{131}I . The different
 F5:3 lines represents the spectra for different concentrations of ^{131}I .

337 spectrum. As a consequence, the minimum detectable activity
 338 of ^{131}I is greater than for other emitters.

339 Identification of ^{131}I was accomplished using several equip-
 340 ment simulations where the filter was impregnated with ^{131}I .
 341 These simulated spectra was combined with an experimental
 342 background spectrum. The resultant spectra is shown in Fig. 5
 343 When the 364.4 keV peak is identified as ^{131}I instead of ^{214}Pb
 344 in the combined spectrum analysed by pGamma, the activity of
 345 ^{131}I ranges between 0.4 Bq/m^3 and 0.8 Bq/m^3 . The latter is its
 346 Minimum Detectable Activity (MDA). For comparison, ^{60}Co
 347 MDA is 0.02402 Bq/m^3 and ^{137}Cs MDA is 0.01662 Bq/m^3
 348 (the air flow for the experimental background spectrum was
 349 $8.870 \text{ m}^3/\text{h}$).

350 B. Determination of Specific Activity

351 The activity concentration of ^{131}I can be determined using
 352 the filter's retention efficiency, which has already been studied
 353 by several authors ([4], [5]). It depends on the iodine species,
 354 particle size and air flow. The equipment for continuous mea-
 355 surement and identification of gamma radioactivity in aerosols
 356 for which this version of the pGamma code has been devel-
 357 oped uses a GF 10 from Hahnemühle fiberglass paper filter.
 358 The fiberglass paper filter can only retain iodine particles. Their
 359 average size can be found in a number of works [6], [7]. Their
 360 diameter typically ranges between $0.2 \mu\text{m}$ and $0.5 \mu\text{m}$. The air
 361 flow of the equipment in normal operation is $110 \text{ m}^3/\text{h}$. In these
 362 conditions, the efficiency of the GF 10 filter for iodide can be
 363 considered as:

$$\varepsilon_f = \begin{cases} 99 - 0\%(\text{for particles}) \\ 0\%(\text{for all the other cases}) \end{cases} \quad (18)$$

364 However, the estimation of filter efficiency is not enough to
 365 determine the activity concentration of ^{131}I . Iodine can appear

as a particle or gas in the form of elemental I or I^2 , organic 366
 (chiefly methyl iodide CH_3I) or inorganic (hypoiodous HOI). 367
 The GF 10 filter does not capture all iodine species (actu- 368
 ally, no filter does); it only retains particles. The only way 369
 to determine total activity concentration in the air is to know 370
 the concentration of iodine particles at the surveillance sta- 371
 tion, but this is impossible because of the special behavior of 372
 iodine transport in the atmosphere. The composition of iodine 373
 can be measured by gas chromatography [8], but this method 374
 is outside the equipment's concept. Several authors have made 375
 estimates from information collected during the Chernobyl or 376
 Fukushima accidents. For example, in [9] it was estimated that 377
 50% of iodine is particles and the rest is gaseous. In his book 378
 [10], A. C. Chamberlain reports some particle-gas rate val- 379
 ues, but they differ significantly from each other. Finally, the 380
 assumption in RASCAL 4.3 code [11], i.e. 33% of iodine par- 381
 ticles, was considered for this work. We must bear in mind that 382
 the final purpose of the equipment is not to perform accurate 383
 measurements of air component concentration, but to provide 384
 information about gamma emitters in the air, such as the above 385
 estimates. 386

VI. ACTION LEVELS 387

Three action levels are provisionally considered²: 388

- 1) Investigation level. 389
- 2) Alert level, according to the concentration limit value in 390
the air for public members (concentration giving by a 391
dose 1 mSv/year). 392
- 3) Alarm level, according to twice the concentration limit 393
value for the public. 394

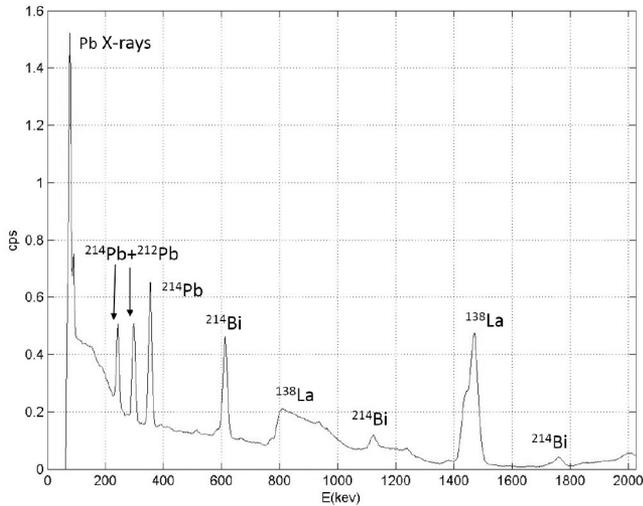
The investigation level is initiated when: 395

- 1) The activity detected exceeds activity concentration 396
thresholds for critical anthropogenic or natural isotopes 397
stored in a library. For the former, the threshold is 398
based on the recommendation of the 18 December 2003 399
Commission, (2004/2/EURATOM) DO L2/36 (6.1.2004). 400
For NORM radionuclides, the threshold is based on sta- 401
tion measurement values (e.g. $2 \cdot 10^{-2} \text{ Bq/m}^3$ for ^{131}I 402
and $3 \cdot 10^{-2} \text{ Bq/m}^3$ for ^{137}Cs). 403
- 2) The analysis report contains one or more unidentified 404
peaks. 405
- 3) The variation of the total number of counts on the detec- 406
tor during consecutive identical periods of time, defined 407
as Relative Variation in the Number of Counts (RVC), 408
exceeds an empirical value. This can be expressed as: 409

$$\frac{dC}{C} = \frac{C(t_2) - C(t_1)}{C(t_1)} \quad (19)$$

Where $C(t_1)$ and $C(t_2)$ are the number of counts on two 410
consecutive identical periods of time. After a year of operation 411
we propose a statistic analysis of the results and a revision of 412
the investigation levels. 413
414

²Provided by the Spanish "Consejo de Seguridad Nuclear" (Nuclear Security Council)



F6:1 Fig. 6. Spectrum analysed, obtained with the equipment for continuous measurement and identification of gamma radioactivity in aerosols with paper filter
 F6:2
 F6:3 ($2'' \times 2''$ LaBr₃ scintillator with digiBASE form ORTEC).

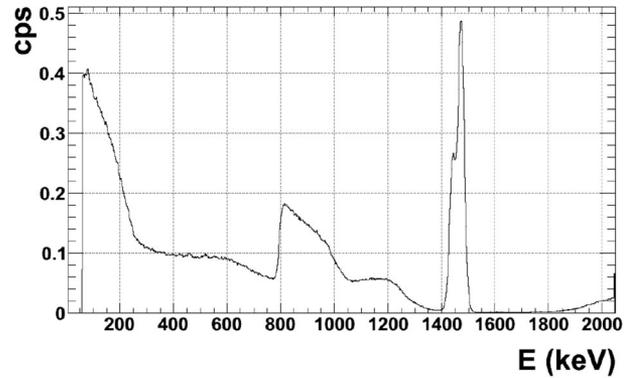
T4:1 TABLE IV
 T4:2 DETECTABLE NORM GAMMA EMITTERS FROM AERSOLS [12]

Emitter	Series	Gammas emitted (keV)
²¹⁴ Pb	²³⁸ U- ²²² Rn	242 (7.2 %), 295.22 (18.5%), 351.93 (35.6 %)
²¹⁴ Bi	²³⁸ U- ²²² Rn	609.31 (45.49 %), 1120.29 (14.09 %), 1764.49 (15.28 %), 1847.42 (2 %), 2204.21 (4.9 %)
²¹² Pb	²³² Th- ²²⁰ Rn	238.63 (43.6 %), 300.09 (3.18 %)
²¹² Bi	²³² Th- ²²⁰ Rn	727.33 (6.74 %), 1620.74 (1.51 %)
²⁰⁸ Tl	²³² Th- ²²⁰ Rn	510.7 (6.29 %), 583.19 (30.6 %), 2614.51 (35.85 %)

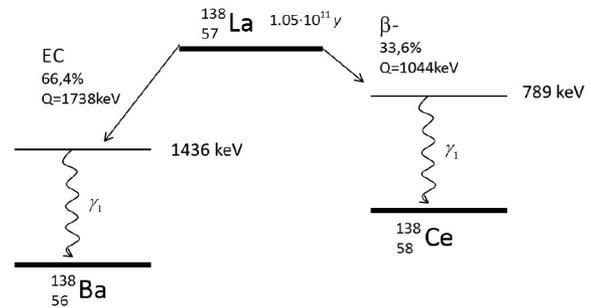
415 VII. ANALYSIS OF A SPECTRUM BY pGAMMA

416 This section presents a pGamma analysis example of a typical
 417 24 h spectrum obtained at Ascó station on 10 January 2015.
 418 The average air flow was 8.35 Bq/m³, there was no rainfall and
 419 the temperature ranged between 7°C and 15°C. The spectrum
 420 is shown in Fig. 6. Note that the peaks caused by NORM, i.e.
 421 detectable gamma emitters of the ²³⁸U series and ²¹²Pb of the
 422 ²³²Th series, appear combined with the intrinsic spectrum of
 423 the LaBr₃ detector. No anthropogenic emitters are observed in
 424 the air.

425 Considering the spectrum in Fig. 6, the components are:
 426 1) Aerosols (mainly NORM) retained on the filter; 2) Cosmic
 427 particles and their interaction with elements on the detection
 428 zone and 3) Intrinsic spectrum of the detector. Table IV illustrates
 429 the detectable gamma peaks of the ²³⁸U and ²³²Th series.
 430 Fig. 7 shows the background spectrum (without air suction or
 431 aerosol retention) generated by the equipment. The cosmic
 432 component is only important at low energies. The level of
 433 intrinsic radiation from the detector is very significant in the
 434 range of interest, even higher than for the component resulting
 435 from aerosols retained on the filter. The radiation generated
 436 in the LaBr₃ crystal is mainly due to ¹³⁸La (Fig. 8). The
 437 most important peaks of the intrinsic spectrum of the LaBr₃



F7:1 Fig. 7. Background spectrum of the equipment of the measurement of gamma
 F7:2 radiation on aerosols by paper filter.



F8:1 Fig. 8. Decay scheme for ¹³⁸La [13].

438 scintillation detector are: 1) between (0-300) keV β continuous;
 439 2) between (700-1100) keV β continuous plus gammas of
 440 789 keV; 3) between (1400-1500) keV gammas of 1436 keV
 441 and a sum peak of gammas of 1436 keV and X-rays of 32 keV.

442 Fig. 9 shows the pGamma report of the analysed spectrum.
 443 The report has four parts: 1) overall parameters of the spec-
 444 trum; 2) specific activity of identified isotopes (first column for
 445 activity, second for error and third for the minimum detectable
 446 activity); 3) identified peaks; and 4) unidentified peaks.

447 Following the list of expected peaks (Table IV), the code
 448 identified: 1) The most probable ²¹⁴Pb peaks; 2) The most prob-
 449 able ²¹⁴Bi peaks; 3) One ²¹²Pb peak; and 4) The 789 keV and
 450 of 1436 keV γ peaks and the combination of the γ of 1436 keV
 451 with the X-ray of 32 keV of the intrinsic spectrum of the LaBr₃
 452 crystal. ²¹²Bi does not appear in the list of identified peaks
 453 because its gamma rays have a very low emission probability
 454 and its full energy peaks are not significant enough. ²⁰⁸Tl has no
 455 identified peaks. From ²⁰⁸Tl, the 510.7 keV peak is very close
 456 to the annihilation peak and is too small to be distinguished by
 457 a scintillator, the same is true of the 583.19 keV peak. This one
 458 is important but very close to the 609.31 keV peak of ²¹⁴Bi,
 459 which is a very significant one, and therefore the overlapping
 460 peak is assigned to ²¹⁴Bi. The 2614.51 keV peak is out of the
 461 operating range of the equipment. It must be remembered that,
 462 with the resolution of scintillators, the 242 keV and 295.22 keV
 463 peaks of ²¹⁴Pb coincide with the ²¹²Pb peaks. Finally, the peak
 464 6 at 514.1 keV is the annihilation peak.

```

[1]
11640368
10179770
[2]
pb212 SerieTh232Rn220 3.15e-001 3.56e-003 6.98e-003 1
bi214 SerieU238Rn222 1.42e+001 8.05e-002 2.88e-001 7 10 11
12 13 17 18
pb214 SerieU238Rn222 1.49e+001 7.23e-002 1.47e-001 2 3
[3]
pb212 1 121.1 241.7 241.8 4.4 115 133 14.4 135857.7
SerieTh232Rn220
bi214 7 307.0 611.0 609.6 9.6 286 326 20.9 302812.7
SerieU238Rn222
bi214 10 467.6 932.4 936.5 5.4 464 479 18.4 5440.6
SerieU238Rn222
bi214 11 561.5 1121.2 1120.2 9.5 542 574 26.7 44514.1
SerieU238Rn222
bi214 12 618.1 1235.2 1238.1 6.3 612 667 23.1 8134.9
SerieU238Rn222
bi214 13 690.1 1380.9 1375.5 7.1 676 695 19.7 4868.6
SerieU238Rn222
bi214 17 876.3 1759.3 1753.0 17.6 846 894 33.7 51719.2
SerieU238Rn222
bi214 18 921.8 1852.1 1841.1 9.9 902 927 30.1 4166.6
SerieU238Rn222
pb214 2 148.9 296.7 296.9 4.8 142 158 17.3 178632.9
SerieU238Rn222
pb214 3 177.6 353.5 353.4 5.5 168 186 14.1 304242.5
SerieU238Rn222
aniq 6 257.8 512.9 514.1 5.3 252 266 17.0 7653.3 Pic
d'aniquilació
la138 9 405.2 807.1 818.3 19.8 391 438 52.7 172130.3
Intrinsec
la138 14 733.3 1468.4 1468.6 7.8 723 780 21.9 229245.5
Intrinsec
[4]

```

F9:1 Fig. 9. Report with the analysis of pGamma.

465

VIII. CONCLUSIONS

466 The pGamma code is an automatic spectrometric analysis
467 system for environmental radiation monitoring equipment with
468 spectrometric capability. The version presented in this paper
469 was specifically designed for our equipment of continuous mea-
470 surement and identification of gamma radioactivity in aerosols
471 with paper filter. Nevertheless, the code is perfectly adaptable to
472 other equipment of the environmental radiological surveillance
473 network of the local Catalan Government. The code identi-
474 fies gamma emitters in the energy spectrum and determines
475 their specific activity. If an emitter is not identified or activ-
476 ity concentration of any identified emitter exceeds an empirical
477 threshold, an alarm is generated.

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A New Code for Spectrometric Analysis for Environmental Radiological Surveillance on Monitors Focused on Gamma Radioactivity on Aerosols

Alfredo de Blas, *Member, IEEE*, Albert Riego, Roger Garcia, Carlos Tapia, Javier Dies, Juan Toral, Enric Batalla, and Pedro Diaz

Abstract—This paper presents a new code for the analysis of gamma spectra generated by an equipment for continuous measurement of gamma radioactivity in aerosols with paper filter. It is called pGamma and has been developed by the Nuclear Engineering Research Group at the Technical University of Catalonia - Barcelona Tech and by Raditel Serveis i Subministraments Tecnològics, Ltd. The code has been developed to identify the gamma emitters and to determine their activity concentration. It generates alarms depending on the activity of the emitters and elaborates reports. Therefore it includes a library with NORM and artificial emitters of interest. The code is being adapted to the monitors of the Environmental Radiological Surveillance Network of the local Catalan Government in Spain (Generalitat de Catalunya) and is used at three stations of the Network.

Index Terms—Environmental radiation, gamma activity, gamma spectrometry, isotope identification.

I. INTRODUCTION

THE use of monitors with spectrometric capability in environmental radiological surveillance networks provides additional information that complements the ambient dose equivalent $H^*(10)$. Based on this premise, our group (Nuclear Engineering Research Group) at the Technical University of Catalonia - Barcelona Tech and Raditel Serveis i Subministraments Tecnològics, Ltd. have developed an equipment for continuous measurement and identification of gamma radioactivity in aerosols with paper filter. Gamma spectrum analysis allows the identification and determination of activity concentration of radiation sources. Excellent commercial software is available for this purpose (among others, GammaVision from ORTEC and Genie2000 from

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Color versions of one or more of the figures in this paper are available online at <http://ieeexplore.ieee.org>.

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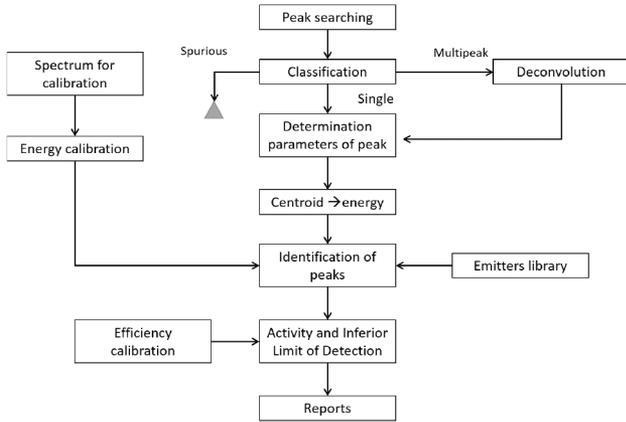
CANBERRA). Nevertheless, it is generic software not intended for automatic analysis since each individual analysis must be conducted by a specialised technician. For this reason, our research group has developed a spectrometric analysis code specifically designed for the above equipment. This code is being used at stations with aerosol monitors on the Environmental Radiological Surveillance Network of the Generalitat de Catalunya (local Catalan Government), Spain¹: Three monitors are currently fully operative: two in the surveillance station of the Ascó and Vandellòs Nuclear Power Plants (both in the province of Tarragona, Spain), and one in the Engineering School of Barcelona, Spain. Two new monitors will shortly be deployed in Roses (province of Girona) and in Puigcerdà (province of Barcelona). The code has been adapted for the analysis of gamma spectra generated by other monitors with spectrometric capability of the Network, i.e. river and direct air monitors. We are starting the tuning of these versions of the code on the monitors for river and direct air.

The code, called pGamma, discriminates spectrum information, identifies emitters appearing on the spectrum, determines activity concentration in Bq/m^3 and generates alarms according to its calculated values. It has a normal operation mode and special operation mode for spectra obtained during energy calibration of the equipment. Additionally, another mode called investigation, is set when the variation in the total number of counts exceeds a certain level, or when a full energy peak is not identified.

The code input is an ASCII text file with the number of counts per channel of the spectrum and a header with information about measurement conditions: date and time of counting initiation and termination, live time, air flow, operation mode, the parameters of the energy calibration curve, etc. Fig. 1 is a general operation diagram of pGamma in normal operation mode.

The paper is organized as follows: Section II describes the treatment of the peaks in the spectrum. In Section III, efficiency calibration is presented. The determination procedure of activity concentration is shown in Section IV. Section V describes the identification of ¹³¹I. In Section VI, the action levels are

¹This network is integrated in the Spanish Nuclear Security Council network.



F1:1 Fig. 1. General diagram of pGamma.

77 presented. An Ascó station spectrum is analysed in Section VII.
78 Finally, conclusions are drawn in Section VIII.

79 II. PEAK TREATMENT

80 The first step in the analysis of the spectrum is the treatment
81 of the peaks: the searching of possible peaks on the spectrum
82 and its identification.

83 A. Peak Search

84 Peak search and analysis is performed after the input spec-
85 trum is smoothed using the Savitzky-Golay method [1]. The
86 smoothing algorithm allows the first, second and third deriva-
87 tives of the spectrum to be obtained. The first derivative is used
88 to determine the local maximums, which are subjected to a sta-
89 tistical test to discriminate real center of peaks (centroids) from
90 spurious transitions. The nearest relative maximums of the first
91 derivative at both sides of the centroid, indicate the channel of
92 the peak boundaries. Using the second and the third derivate,
93 single peaks are discriminated from multiplets. With the cen-
94 troid and the boundaries of the peaks, the net area and the
95 background are determined. The last step is the fitting of each
96 peaks to a Gaussian curve. It allows the determination of the full
97 width at half maximum (FWHM) and perform the chi-square
98 test for further discrimination. With this information, pGamma
99 generates a first list of possible peaks with their characteris-
100 tics. The calculated parameters are centroid, lower and upper
101 boundaries of the possible peak, full width at half maximum,
102 gross area and net area.

103 B. Peak Identification

104 1) *Energy Calibration:* The use of energy calibration
105 makes it possible to state the centroid, FWHM and limits
106 of peaks in energy units instead of number of channels. The
107 relationship between energy and adopted number of channel is:

$$E = a + bC + cC^2 \quad (1)$$

108 Where C is the channel and a , b and c are the parameters of the
109 calibration curve. Energy calibration must be performed prior
110 to installation of equipment in a station.

2) *Peak Identification:* Peaks can be identified using the
111 library of gamma emitters of interest included in the code.
112 This library has been created by us using information from [2].
113 For aerosol equipment, these emitters are the NORM radionu-
114 clides (^{238}U - ^{222}Rn series, ^{232}Th - ^{220}Rn series, ^{40}K and ^7Be)
115 and artificial emitters from nuclear power plant discharges
116 (e.g. radio-iodines, ^{137}Cs , ^{134}Cs and ^{132}Te) or industrial appli-
117 cations. Information about emitters includes energy and the
118 emission probability of emitted gamma rays, empirical thresh-
119 olds for NORM emitters to decide if pGamma must switch to
120 investigation mode and derived concentration limits in the air.
121

The centroid energy of each peak is compared with the
122 energy of gamma emitters. For similar values, the emitters are
123 considered candidates for that peak and a list of candidates
124 is generated. To identify a set of peaks with a radionuclide
125 with multiple gamma emissions, the ratio between intensity
126 of detected gammas and total intensity of the gammas of the
127 emitter in the library is used. The relationship between the
128 members of ^{238}U - ^{222}Rn series and ^{232}Th - ^{220}Rn series is con-
129 sidered too. pGamma cannot determine activity concentration
130 of unidentified peaks.
131

III. DETECTION EFFICIENCY CALIBRATION

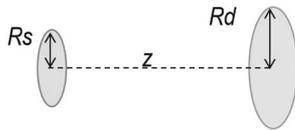
132 In order to determine activity concentration, the code needs
133 information about the detection efficiency. Since the factors
134 integrating the detection efficiency are very difficult to deter-
135 mine due to a complex geometry and a wide range of possible
136 energies, efficiency calibration must be performed. First, the
137 same geometry of the paper filter (source), detector and detec-
138 tion volume of the equipment during normal operation must
139 be reproduced. This implies preparing a calibration source
140 by shaping it into a disc with the same diameter of aerosols
141 retained on the filter, and the same self-absorption (negligi-
142 ble) and gamma emitters of several energies in the range from
143 100 keV to 2000 keV. In consequence, the analytical determi-
144 nation of efficiency is cumbersome. The solution consists in
145 using a calibration source with similar characteristics to the
146 source to be measured and determining efficiency from experi-
147 mental values. To ensure unbiased calibration, the calibration
148 source should be identical to the radioactive samples in all
149 aspects affecting detection [3]. If no calibration source is avail-
150 able or more points are required for adjusting the efficiency
151 curve, Monte Carlo n-Particle (MCNP5, version 2.6.0) provide
152 a further possibility for calculating the detector efficiency.
153

We used simulated ^{131}I and commercial ^{60}Co sources
154 (Table I) with similar geometry and materials to those of the
155 impregnated filter of the monitor. The simulated ^{131}I source
156 is composed of two radioisotopes: ^{133}Ba and ^{137}Cs . The com-
157 bination of both emitters generates a gamma energy spectrum
158 whose full energy peaks have centroids and a relationship of
159 areas very similar to the ^{131}I spectrum. The simulated ^{131}I
160 source has the same area as the impregnated filter. The emitter
161 material is deposited on a polymeric membrane with a stainless
162 steel backing of 0.762 mm thickness and an 0.9 mg/cm^2 alu-
163 minized mylar window. Gamma absorption of the window and
164 the backscattering on the back of the membrane are negligible.
165 Therefore, it can be considered that the material and geometry
166

T1:1
T1:2
T1:3

TABLE I
CHARACTERISTICS OF THE CALIBRATION SOURCES (REFERENCE FOR
ACTIVITY: 01-JANUARY-2014)

(a)	
Nuclides	^{137}Cs , ^{133}Ba
^{137}Cs Activity	0.4096 kBq
^{133}Ba Activity	4.222 kBq
^{131}I Activity	3.684 kBq
Manufacturer	Eckert & Ziegeler
Model	EAB-131(s)-PL
Source No.	K9-928
(b)	
Activity	3.863 kBq
Manufacturer	Eckert & Ziegeler
Model	EAB-060-PL
Source No.	K9-925



F2:1 Fig. 2. Geometry for equations (2) and (4).

167 of the source have the same behaviour as the impregnated paper
168 filter. To simulate the detector response on the MCNP5 sim-
169 ulation we had used the FWHM obtained from the peaks on
170 the spectra for the calibration sources (^{131}I and ^{60}Co) with the
171 experimental setup. The function used is:

$$FWHM = 0.008 + 0.013 \cdot \sqrt{E + 0.593 \cdot E^2} \quad (2)$$

172 where E is the energy in MeV and the FWHM obtained is also
173 in MeV.

174 Experimental efficiencies were obtained by analyzing the
175 experimental spectra of ^{131}I and ^{60}Co by pGamma. Before the
176 calculation of the efficiency, the net counts of the experimental
177 peaks had to be compensated by a correction factor. The source
178 capsule elevated the membrane where the radioactive material
179 is deposited such that the membrane is slightly higher than the
180 paper filter (about 0.283 cm). Assuming that both sources, the
181 impregnated paper filter in normal operation and the calibration
182 source, are discs and the detector window is a disc too (geome-
183 try on Fig. 2), we could determine the geometric factor in both
184 cases using the following expansion in power series:

$$gf_i = \frac{1}{2} \left(1 - \frac{1}{R} \right) - \frac{3 \cdot \omega^2 \psi^2}{16 \cdot R^5} \cdot \left[1 - \frac{5 \psi^2}{6 \cdot R^4} \cdot \left(1 - \frac{3 \cdot \omega^2}{4} \right) + \frac{35 \cdot \psi^4}{48 \cdot R^8} \left(1 - \frac{5 \cdot \omega^2}{2} + \frac{5 \cdot \omega^4}{8} \right) + \dots \right] \quad (3)$$

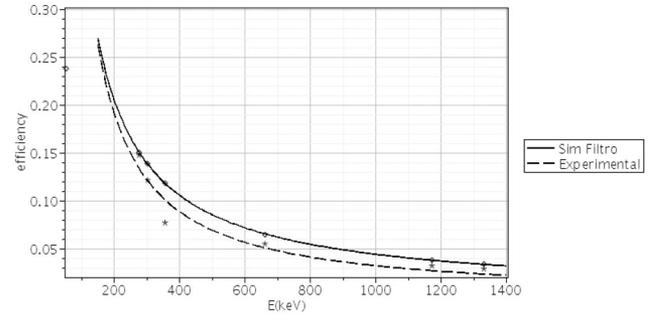
185 the factors R , ψ and ω are determined as:

$$\psi = \frac{R_S}{z}; \omega = \frac{R_d}{z}; R = \sqrt{1 + \omega^2} \quad (4)$$

186 where R_S is the source radius, R_d is the detector radius, z is the
187 distance between the source and the detector. Equation (4) can

TABLE II
EXPERIMENTAL DETECTION EFFICIENCY

Peak	E_i (keV)	N_i (cps)	S_i (ph/s)	ε_d
^{131}I sim	53.160	57.990	88.510	0.70319
^{131}I sim	276.390	40.801	296.150	0.14786
^{131}I sim	303.000	86.024	758.550	0.12170
^{131}I sim	356.010	184.860	2566.500	0.017730
^{131}I sim	661.500	19.830	385.100	0.05527
^{60}Co	1173.000	108.028	3573.290	0.03244
^{60}Co	1332.000	97.690	3579.020	0.02929



F3:1 Fig. 3. Curve of efficiency.

be applied either for the calibration source or the impregnated
filter. Then the correction factor cf is:

$$cf = \frac{gf_{calibration\ source}}{gf_{filter}} \quad (5)$$

where gf is the geometric factor. With this method, the cor-
rection factor is 0.93680. By using seven MCNP5 code sim-
ulations, one for each energy, the mean correction factor is
0.93184. Efficiency ε_d for energy E_i is determined as:

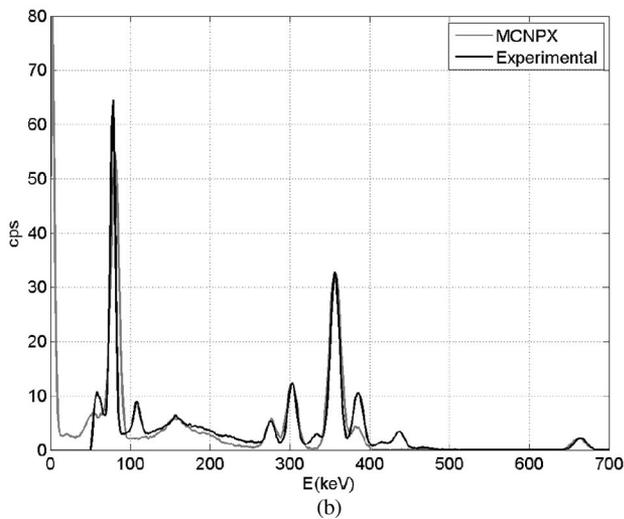
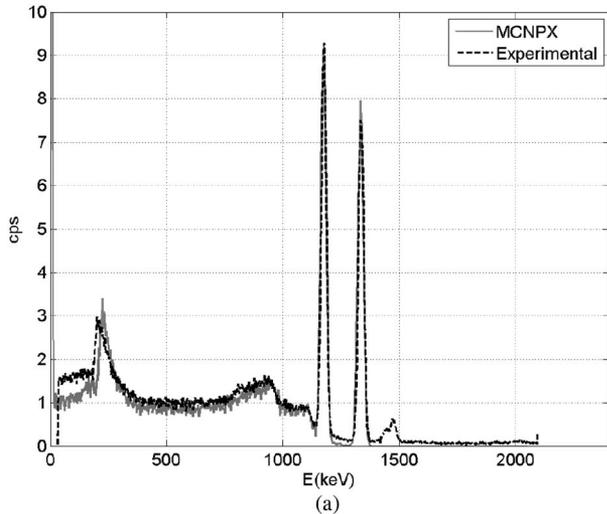
$$\varepsilon_d = \frac{1}{cf} \cdot \frac{N_i}{S_i} \quad (6)$$

where N_i represents the net count rate of the full energy peak
corresponding to the source gamma rays with energy E_i and S_i
is the intensity of photons with energy E_i . Intensity is deter-
mined for the two components of the simulated ^{131}I : ^{137}Cs and
 ^{133}Ba , which are treated as two separate sources.

$$S_i = \nu_i \cdot A_i(0) e^{-T_t \cdot \lambda_i}; i = ^{137}\text{Cs}, ^{133}\text{Ba} \quad (7)$$

where ν_i is the emission probability of photons with energy E_i ,
 $A_i(0)$ is the activity on the calibration date of the source, T_t is
the time between the calibration date of the source and the day
of the experiment and λ_i is the disintegration constant. With
equations (6) and (7) we find the values of ε_d on Table II. A
second efficiency curve, obtained from MCNP5, is compared
with that obtained from experimental values, Fig. 3. As can
be seen, the two curves are very close, with the exception of
experimental efficiency for 380 keV. By fitting the experimen-
tal points without considering the 380 keV experimental value,
efficiency is:

$$\varepsilon_d = \begin{cases} 0.43792 & 100\text{keV} \leq E \leq 200\text{keV} \\ 60.71703 \cdot E^{-1.07096} & E > 200\text{keV} \end{cases} \quad (8)$$



F4:1 Fig. 4. Comparison between simulated and experimental spectra. Experimental
 F4:2 setup: $\text{LaBr}_3 2'' \times 2''$ scintillator and the DigiBASE compact instrumentation
 F4:3 from ORTEC. The source is inside the volume of detection equipment for con-
 F4:4 tinuous measurement and identification of gamma radioactivity in aerosols with
 F4:5 paper filter. (a) Co-60 (b) Simulated I-131.

211 Since Monte Carlo methods were used to complement the
 212 experimental values to obtain the efficiency curve, the equip-
 213 ment simulation model must be benchmarked. In fact, without
 214 the geometry correction, Fig. 3 itself could be a validation of
 215 the Monte Carlo simulations. It is interesting to analyze and
 216 compare the experimental and Monte Carlo simulated spec-
 217 tra of the equipment with the calibration source instead of the
 218 paper filter. As the sources were very close to the front win-
 219 dows of the detector, some true coincidence summing peaks
 220 appeared on the spectrum. These cannot be reproduced with
 221 a single MCNP5 simulation, but as we want to compare the
 222 procedure used to obtain the efficiency curve and these sum
 223 peaks are not required for the efficiency calculations, we will
 224 not simulate them. Fig. 4 compares simulated and experimen-
 225 tal ^{60}Co and simulated source of ^{131}I spectra. The simulation
 226 reproduces perfectly the processes that not depend on the
 227 electronic instrumentation modules. As can be seen, backscat-
 228 tering peaks, Compton continuum, Compton valley and two

TABLE III
 COMPARISON OF NET COUNTS OF PEAKS

		Net Counts ^{60}Co (cps)			
E (keV)	Experim.	MCNP			
1173	107.95	122.13			
1332	97.62	112.34			
		Net Cnts ^{131}I (cps)		Gross counts ^{131}I (cps)	
E (keV)	Experim.	MCNPX	Experim.	MCNP	
60	24.11	8.60	57.99	65.19	
80	189.18	303.36	260.91	380.45	
276	21.02	17.61	40.80	44.48	
303	63.31	73.27	86.02	104.65	
356	184.86	249.07	231.13	291.88	
661.5	19.83	22.51	22.32	22.79	

229 full energy peaks appear. The differences between both spectra
 230 are due to the fact that background radiation is not simulated.
 231 Intrinsic radiation from the $\text{LaBr}_3(\text{Ce})$ scintillator, i.e. ^{138}La
 232 and ^{138}La plus X-rays peaks at 1468 keV, appears in the experi-
 233 mental spectrum. This component was not simulated, but is not
 234 relevant for efficiency calibration. Moreover, at low energies
 235 the simulated spectrum has fewer counts than the experimental
 236 one. This is because the Monte Carlo model does not simulate
 237 the background radiation components of cosmic rays or earth's
 238 NORMs, resulting in fewer X-rays on the lead shielding (peak
 239 near 80 keV in the spectrum). In our model, only X-rays from
 240 absorption of the calibration source radiation are generated.
 241 Again, this is not important for efficiency calculation.

242 Since the simulated ^{131}I has many peaks at low energies, the
 243 effect of sum peaks is more noticeable, at least in our energy
 244 range (0-2048 keV). Monte Carlo does not generate sum peaks,
 245 as can be seen in Fig. 4. For full energy peaks with little or no
 246 influence of sum peaks, simulation results are good. Table III
 247 compares the net areas of full energy peaks of simulated and
 248 experimental spectra. Peaks influenced by a sum peak are not
 249 used in efficiency calibration.

IV. DETERMINATION OF THE ACTIVITY CONCENTRATION 250

251 After identification of gamma emitters in the spectrum, the
 252 spectrometric analysis system determines their specific activity
 253 in the air \tilde{A}_a , expressed in Bq/m^3 , according to the net area of
 254 the most efficient emitter of the isotope. However, the obtained
 255 spectrum corresponds to the number of counts caused by the
 256 radionuclide concentration on the filter, which is different from
 257 the radionuclide concentration in the air. Because the activity on
 258 the filter A_f is proportional to the number of captured atoms,
 259 a balance of activity can be performed in the same way as for
 260 concentration. The radionuclide concentration of an analysed
 261 emitter on the filter at a certain time depends on three factors:
 262 1) the number of captured atoms of this emitter; 2) decays of
 263 the parent nuclei of this emitter, and 3) decays of this emitter.

264 Activity on the filter A_f is proportional to activity in the air
 265 \tilde{A}_a , air flow Q , and retention efficiency of the filter ε_f [4], [5].
 266 Nevertheless, as activity on the filter implies a balance of matter
 267 on the filter, concentration of each radionuclide depends on its

268 position in the decay series. In the case of aerosol equipment,
269 the radionuclides are belonging to the ^{238}U - ^{222}Rn series, the
270 ^{232}Th - ^{220}Rn (chiefly ^{212}Pb , ^{212}Bi and ^{208}Tl) series or to none
271 of the series. A slightly different statement is required for each
272 case.

273 A. Radionuclides not Belonging to a Series

274 Most artificial emitters fall in this category. The balance of
275 matter in this case is simply:

$$\frac{dA_f}{dt} = Q\varepsilon_f \tilde{A}_a(0) - \lambda_i A_f \quad (9)$$

276 Equation (9) can be solved by integration and as result the filter
277 activity of the emitter analysed is:

$$A_f(t) = \frac{Q \cdot \varepsilon_f \cdot \tilde{A}_a}{\lambda} \int_0^T 1 - e^{-\lambda t} dt \quad (10)$$

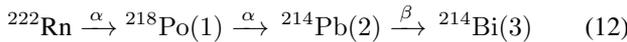
278 Then, from Equation (10), the specific activity \tilde{A}_a in the air for
279 emitter i (expressed Bq/m³) is:

$$\tilde{A}_a = \frac{\lambda_i N_i T}{Q \varepsilon_f \varepsilon_d \nu_i \int_0^T (1 - e^{-\lambda_i t}) dt} \quad (11)$$

280 Where λ_i is the decay constant for radionuclide i , N_i the net
281 area (counts per second) of the full energy peak of the most
282 probable gamma of radionuclide i , T the detection time (live
283 time fo the Multiple Channel Analyzer), Q the average air-
284 flow, ε_f the filter efficiency, ε_d the detection efficiency for the
285 most probable gamma of radionuclide i , and ν_i the emission
286 probability of the most probable gamma of radionuclide i .

287 B. ^{238}U - ^{222}Rn Series

288 The gamma emitters in the air of interest belonging to the
289 decay series of ^{238}U are the descendents of ^{222}Rn :



290 Numbers in parenthesis in the chain (12) represent the nomen-
291 clature used for the radionuclides in the following equations.
292 These radionuclides are present on all countings because they
293 come from the earth and surrounding buildings. Among them
294 we can find some gamma emitters of certain importance, mainly
295 ^{214}Pb and ^{214}Bi . The equations of balance of matter for chain
296 (12) are:

$$\begin{aligned} \frac{dA_{f1}}{dt} &= \tilde{A}_{a1}(0)Q\varepsilon_f - \lambda_1 A_{f1} \\ \frac{dA_{f2}}{dt} &= \lambda_1 A_{f1} + \tilde{A}_{a2}(0)Q\varepsilon_f - \lambda_2 A_{f2} \\ \frac{dA_{f3}}{dt} &= \lambda_2 A_{f2} + \tilde{A}_{a3}(0)Q\varepsilon_d - \lambda_3 A_{f3} \end{aligned} \quad (13)$$

297 ^{218}Po reaches its equilibrium at 95% in 13 minutes. By com-
298 paring the partial time of detection of 1 h and the total time of
299 detection of 24 h, the hypothesis that ^{218}Po and ^{214}Pb are in
300 equilibrium in the air can be supported.

$$\tilde{A}_{a1}(0) = \tilde{A}_{a2}(0) \quad (14)$$

By considering the equilibrium condition Eq. (14) and devel- 301
oping the balance equations Eq. (13), the specific activities of 302
 ^{214}Pb and ^{214}Bi can be determined by: 303

$$\tilde{A}_{a2}(0) = \frac{1}{2} \frac{N_2 T \lambda_2}{Q \nu_2 \varepsilon_f \varepsilon_{d2} \int_0^T (1 - e^{-\lambda_2 t}) dt} \quad (15)$$

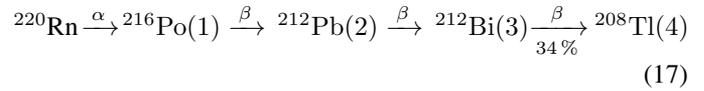
The equation of balance of matter for chain 12 is (for simplicity 304
we omit the explicit time dependences here): 305

$$\begin{aligned} \tilde{A}_{a3}(0) &= \frac{\lambda_3 N_3 T}{Q \nu_3 \varepsilon_f \varepsilon_{d3} \int_0^T (1 - e^{-\lambda_3 t}) dt} \\ &- \frac{2 \tilde{A}_{a2}(0)}{\int_0^T (1 - e^{-\lambda_3 t}) dt} \left[\int_0^T (1 - e^{-\lambda_3 t}) dt \right. \\ &\left. + \frac{\lambda_3}{\lambda_2 - \lambda_3} \int_0^T (e^{-\lambda_2 t} - e^{-\lambda_3 t}) dt \right] \end{aligned} \quad (16)$$

Emission probability ν_i , the net area of the full energy peak 306
 N_i and detection efficiency ε_i correspond to the most probable 307
gamma of the analysed radionuclide. 308

309 C. ^{232}Th - ^{220}Rn Series

As in the previous case, only the nuclides at the end of the 310
chain are of interest in this series, i.e. from ^{220}Rn : 311

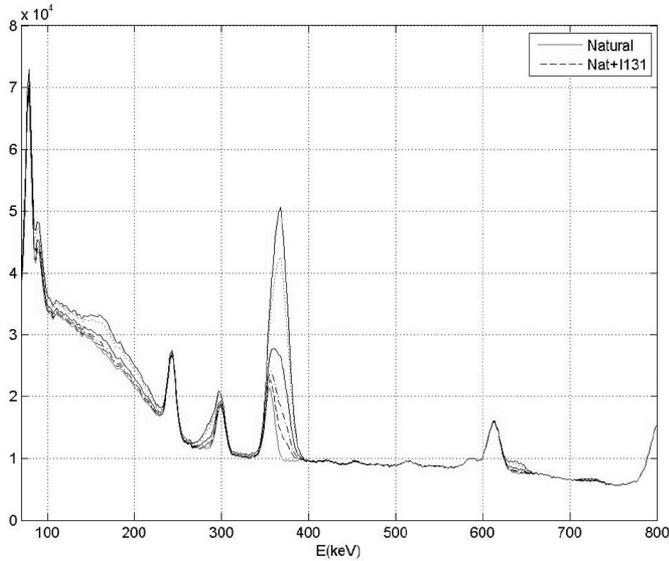


These radionuclides are always present in air aspiration, like 312
in the previous series. Among them, ^{212}Pb , ^{212}Bi and ^{208}Tl 313
are detected by the equipment. ^{212}Pb bifurcates to ^{208}Tl with 314
35.94% and to ^{212}Po with 64.06%, but ^{212}Po is not an impor- 315
tant γ emitter. However, in the location areas of our monitoring 316
stations there are fewer of those than of the ^{238}U - ^{222}Rn series, 317
and in typical spectra generated by the equipment only the pres- 318
ence of ^{212}Pb is detected. The balance of matter is similar to 319
that of the previous series Eq. 13. The half-life of ^{216}Po is 320
0.15 s. Hence, it can be assumed that at the end of detection 321
equilibrium exists between ^{216}Po and ^{212}Bi . 322

323 V. ANALYSIS OF ^{131}I

324 A. Identification and Minimum Detectable Activity (MDA)

Full energy peaks of ^{131}I gamma emitters are very close 325
to some full energy peaks of ^{212}Pb , ^{214}Pb and ^{214}Bi . These 326
natural emitters belong to the ^{222}Rn radon and ^{220}Rn thoron 327
series, and are always present in the spectrum. Overlapping 328
of iodine gamma lines with other emitters occurs between 329
the 284.2 keV, 364.4 keV and 636.97 keV peaks of ^{131}I . 330
The 722.89 keV peak does not overlap with those of natu- 331
ral emitters, but it has very low emission probability. Like the 332
80.18 keV of ^{131}I , that does not overlap with other peaks, but is 333
outside the linear operating range of the scintillator. Therefore, 334
to identify ^{131}I we can only use full energy peaks overlapped 335
with natural emitter peaks, which are always present in the 336



F5:1 Fig. 5. Spectra of ^{131}I . Obtained by a combination of experimental back-
 F5:2 ground and simulations with the filter impregnated with ^{131}I . The different
 F5:3 lines represents the spectra for different concentrations of ^{131}I .

337 spectrum. As a consequence, the minimum detectable activity
 338 of ^{131}I is greater than for other emitters.

339 Identification of ^{131}I was accomplished using several equip-
 340 ment simulations where the filter was impregnated with ^{131}I .
 341 These simulated spectra was combined with an experimental
 342 background spectrum. The resultant spectra is shown in Fig. 5
 343 When the 364.4 keV peak is identified as ^{131}I instead of ^{214}Pb
 344 in the combined spectrum analysed by pGamma, the activity of
 345 ^{131}I ranges between 0.4 Bq/m^3 and 0.8 Bq/m^3 . The latter is its
 346 Minimum Detectable Activity (MDA). For comparison, ^{60}Co
 347 MDA is 0.02402 Bq/m^3 and ^{137}Cs MDA is 0.01662 Bq/m^3
 348 (the air flow for the experimental background spectrum was
 349 $8.870 \text{ m}^3/\text{h}$).

350 B. Determination of Specific Activity

351 The activity concentration of ^{131}I can be determined using
 352 the filter's retention efficiency, which has already been studied
 353 by several authors ([4], [5]). It depends on the iodine species,
 354 particle size and air flow. The equipment for continuous mea-
 355 surement and identification of gamma radioactivity in aerosols
 356 for which this version of the pGamma code has been devel-
 357 oped uses a GF 10 from Hahnemühle fiberglass paper filter.
 358 The fiberglass paper filter can only retain iodine particles. Their
 359 average size can be found in a number of works [6], [7]. Their
 360 diameter typically ranges between $0.2 \mu\text{m}$ and $0.5 \mu\text{m}$. The air
 361 flow of the equipment in normal operation is $110 \text{ m}^3/\text{h}$. In these
 362 conditions, the efficiency of the GF 10 filter for iodide can be
 363 considered as:

$$\varepsilon_f = \begin{cases} 99 - 0\%(\text{for particles}) \\ 0\%(\text{for all the other cases}) \end{cases} \quad (18)$$

364 However, the estimation of filter efficiency is not enough to
 365 determine the activity concentration of ^{131}I . Iodine can appear

as a particle or gas in the form of elemental I or I^2 , organic 366
 (chiefly methyl iodide CH_3I) or inorganic (hypoiodous HOI). 367
 The GF 10 filter does not capture all iodine species (actu- 368
 ally, no filter does); it only retains particles. The only way 369
 to determine total activity concentration in the air is to know 370
 the concentration of iodine particles at the surveillance sta- 371
 tion, but this is impossible because of the special behavior of 372
 iodine transport in the atmosphere. The composition of iodine 373
 can be measured by gas chromatography [8], but this method 374
 is outside the equipment's concept. Several authors have made 375
 estimates from information collected during the Chernobyl or 376
 Fukushima accidents. For example, in [9] it was estimated that 377
 50% of iodine is particles and the rest is gaseous. In his book 378
 [10], A. C. Chamberlain reports some particle-gas rate val- 379
 ues, but they differ significantly from each other. Finally, the 380
 assumption in RASCAL 4.3 code [11], i.e. 33% of iodine par- 381
 ticles, was considered for this work. We must bear in mind that 382
 the final purpose of the equipment is not to perform accurate 383
 measurements of air component concentration, but to provide 384
 information about gamma emitters in the air, such as the above 385
 estimates. 386

387 VI. ACTION LEVELS

Three action levels are provisionally considered²: 388

- 1) Investigation level. 389
- 2) Alert level, according to the concentration limit value in 390
the air for public members (concentration giving by a 391
dose 1 mSv/year). 392
- 3) Alarm level, according to twice the concentration limit 393
value for the public. 394

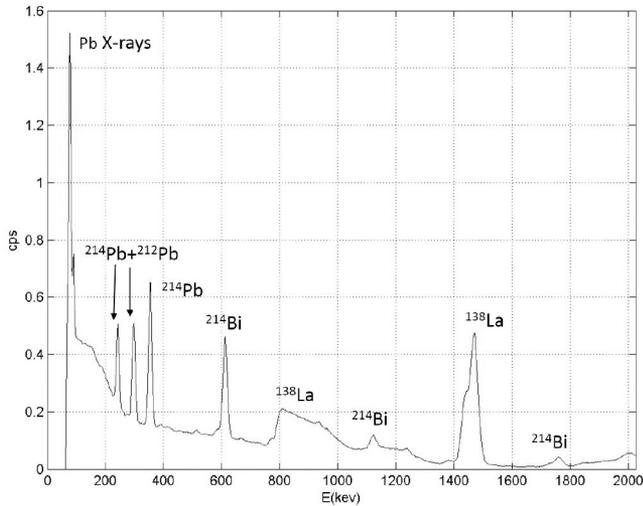
The investigation level is initiated when: 395

- 1) The activity detected exceeds activity concentration 396
thresholds for critical anthropogenic or natural isotopes 397
stored in a library. For the former, the threshold is 398
based on the recommendation of the 18 December 2003 399
Commission, (2004/2/EURATOM) DO L2/36 (6.1.2004). 400
For NORM radionuclides, the threshold is based on sta- 401
tion measurement values (e.g. $2 \cdot 10^{-2} \text{ Bq/m}^3$ for ^{131}I 402
and $3 \cdot 10^{-2} \text{ Bq/m}^3$ for ^{137}Cs). 403
- 2) The analysis report contains one or more unidentified 404
peaks. 405
- 3) The variation of the total number of counts on the detec- 406
tor during consecutive identical periods of time, defined 407
as Relative Variation in the Number of Counts (RVC), 408
exceeds an empirical value. This can be expressed as: 409

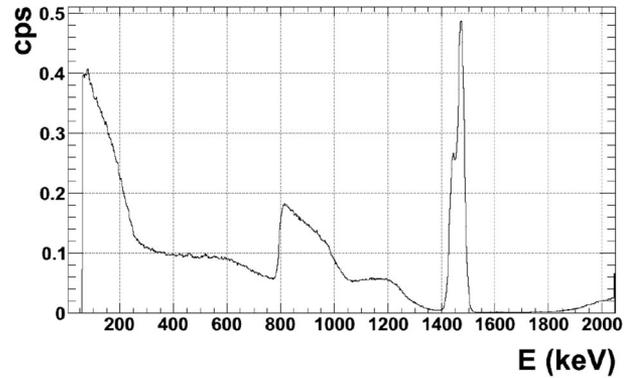
$$\frac{dC}{C} = \frac{C(t_2) - C(t_1)}{C(t_1)} \quad (19)$$

410
 411 Where $C(t_1)$ and $C(t_2)$ are the number of counts on two
 412 consecutive identical periods of time. After a year of operation
 413 we propose a statistic analysis of the results and a revision of
 414 the investigation levels.

²Provided by the Spanish "Consejo de Seguridad Nuclear" (Nuclear Security Council)



F6:1 Fig. 6. Spectrum analysed, obtained with the equipment for continuous measurement and identification of gamma radioactivity in aerosols with paper filter
 F6:2
 F6:3 ($2'' \times 2''$ LaBr₃ scintillator with digiBASE form ORTEC).



F7:1 Fig. 7. Background spectrum of the equipment of the measurement of gamma
 F7:2 radiation on aerosols by paper filter.

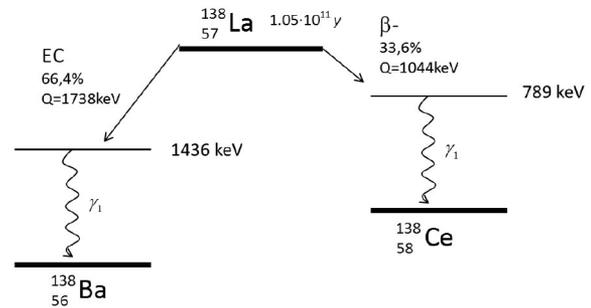
T4:1 TABLE IV
 T4:2 DETECTABLE NORM GAMMA EMITTERS FROM AERSOLS [12]

Emitter	Series	Gammas emitted (keV)
²¹⁴ Pb	²³⁸ U- ²²² Rn	242 (7.2 %), 295.22 (18.5%), 351.93 (35.6 %)
²¹⁴ Bi	²³⁸ U- ²²² Rn	609.31 (45.49 %), 1120.29 (14.09 %), 1764.49 (15.28 %), 1847.42 (2 %) 2204.21 (4.9 %)
²¹² Pb	²³² Th- ²²⁰ Rn	238.63 (43.6 %), 300.09 (3.18 %)
²¹² Bi	²³² Th- ²²⁰ Rn	727.33 (6.74 %), 1620.74 (1.51 %)
²⁰⁸ Tl	²³² Th- ²²⁰ Rn	510.7 (6.29 %), 583.19 (30.6 %), 2614.51 (35.85 %)

415 VII. ANALYSIS OF A SPECTRUM BY pGAMMA

416 This section presents a pGamma analysis example of a typical
 417 24 h spectrum obtained at Ascó station on 10 January 2015.
 418 The average air flow was 8.35 Bq/m³, there was no rainfall and
 419 the temperature ranged between 7°C and 15°C. The spectrum
 420 is shown in Fig. 6. Note that the peaks caused by NORM, i.e.
 421 detectable gamma emitters of the ²³⁸U series and ²¹²Pb of the
 422 ²³²Th series, appear combined with the intrinsic spectrum of
 423 the LaBr₃ detector. No anthropogenic emitters are observed in
 424 the air.

425 Considering the spectrum in Fig. 6, the components are:
 426 1) Aerosols (mainly NORM) retained on the filter; 2) Cosmic
 427 particles and their interaction with elements on the detection
 428 zone and 3) Intrinsic spectrum of the detector. Table IV illustrates
 429 the detectable gamma peaks of the ²³⁸U and ²³²Th series.
 430 Fig. 7 shows the background spectrum (without air suction
 431 or aerosol retention) generated by the equipment. The cosmic
 432 component is only important at low energies. The level of
 433 intrinsic radiation from the detector is very significant in the
 434 range of interest, even higher than for the component resulting
 435 from aerosols retained on the filter. The radiation generated
 436 in the LaBr₃ crystal is mainly due to ¹³⁸La (Fig. 8). The
 437 most important peaks of the intrinsic spectrum of the LaBr₃



F8:1 Fig. 8. Decay scheme for ¹³⁸La [13].
 F8:2

scintillation detector are: 1) between (0-300) keV β continuous; 438
 2) between (700-1100) keV β continuous plus gammas of 439
 789 keV; 3) between (1400-1500) keV gammas of 1436 keV 440
 and a sum peak of gammas of 1436 keV and X-rays of 32 keV. 441

Fig. 9 shows the pGamma report of the analysed spectrum. 442
 The report has four parts: 1) overall parameters of the spec- 443
 trum; 2) specific activity of identified isotopes (first column for 444
 activity, second for error and third for the minimum detectable 445
 activity); 3) identified peaks; and 4) unidentified peaks. 446

Following the list of expected peaks (Table IV), the code 447
 identified: 1) The most probable ²¹⁴Pb peaks; 2) The most prob- 448
 able ²¹⁴Bi peaks; 3) One ²¹²Pb peak; and 4) The 789 keV and 449
 of 1436 keV γ peaks and the combination of the γ of 1436 keV 450
 with the X-ray of 32 keV of the intrinsic spectrum of the LaBr₃ 451
 crystal. ²¹²Bi does not appear in the list of identified peaks 452
 because its gamma rays have a very low emission probability 453
 and its full energy peaks are not significant enough. ²⁰⁸Tl has no 454
 identified peaks. From ²⁰⁸Tl, the 510.7 keV peak is very close 455
 to the annihilation peak and is too small to be distinguished by 456
 a scintillator, the same is true of the 583.19 keV peak. This one 457
 is important but very close to the 609.31 keV peak of ²¹⁴Bi, 458
 which is a very significant one, and therefore the overlapping 459
 peak is assigned to ²¹⁴Bi. The 2614.51 keV peak is out of the 460
 operating range of the equipment. It must be remembered that, 461
 with the resolution of scintillators, the 242 keV and 295.22 keV 462
 peaks of ²¹⁴Pb coincide with the ²¹²Pb peaks. Finally, the peak 463
 6 at 514.1 keV is the annihilation peak. 464

```

[1]
11640368
10179770
[2]
pb212 SerieTh232Rn220 3.15e-001 3.56e-003 6.98e-003 1
bi214 SerieU238Rn222 1.42e+001 8.05e-002 2.88e-001 7 10 11
12 13 17 18
pb214 SerieU238Rn222 1.49e+001 7.23e-002 1.47e-001 2 3
[3]
pb212 1 121.1 241.7 241.8 4.4 115 133 14.4 135857.7
SerieTh232Rn220
bi214 7 307.0 611.0 609.6 9.6 286 326 20.9 302812.7
SerieU238Rn222
bi214 10 467.6 932.4 936.5 5.4 464 479 18.4 5440.6
SerieU238Rn222
bi214 11 561.5 1121.2 1120.2 9.5 542 574 26.7 44514.1
SerieU238Rn222
bi214 12 618.1 1235.2 1238.1 6.3 612 667 23.1 8134.9
SerieU238Rn222
bi214 13 690.1 1380.9 1375.5 7.1 676 695 19.7 4868.6
SerieU238Rn222
bi214 17 876.3 1759.3 1753.0 17.6 846 894 33.7 51719.2
SerieU238Rn222
pb214 2 148.9 296.7 296.9 4.8 142 158 17.3 178632.9
SerieU238Rn222
pb214 3 177.6 353.5 353.4 5.5 168 186 14.1 304242.5
SerieU238Rn222
aniq 6 257.8 512.9 514.1 5.3 252 266 17.0 7653.3 Pic
d'aniquilació
la138 9 405.2 807.1 818.3 19.8 391 438 52.7 172130.3
Intrinsec
la138 14 733.3 1468.4 1468.6 7.8 723 780 21.9 229245.5
Intrinsec
[4]

```

F9:1 Fig. 9. Report with the analysis of pGamma.

465

VIII. CONCLUSIONS

466 The pGamma code is an automatic spectrometric analysis
467 system for environmental radiation monitoring equipment with
468 spectrometric capability. The version presented in this paper
469 was specifically designed for our equipment of continuous mea-
470 surement and identification of gamma radioactivity in aerosols
471 with paper filter. Nevertheless, the code is perfectly adaptable to
472 other equipment of the environmental radiological surveillance
473 network of the local Catalan Government. The code identi-
474 fies gamma emitters in the energy spectrum and determines
475 their specific activity. If an emitter is not identified or activ-
476 ity concentration of any identified emitter exceeds an empirical
477 threshold, an alarm is generated.

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