

Intrapulse x-ray parametric amplification in high-order-harmonic generation

Carles Serrat*

Universitat Politècnica de Catalunya, Colom 11, 08222 Terrassa, Barcelona, Spain

(Received 20 April 2016; published xxxxxx)

We demonstrate strong-field-driven impulsive XUV-x-ray parametric amplification (IXPA) processes in high-order harmonic generation at the single-atom level by using *ab initio* calculations. We consider the example of Li^+ ions exposed simultaneously to an intense IR pulse and a weak 200-as XUV-x-ray pulse with central photon energies varying from 90 to 400 eV. We determine optimal parameter ranges and the precise delays between the IR and the XUV-x-ray pulses for IXPA to occur. The present results might be a guide to achieve exponential growth of the XUV-x-ray signal in tabletop XUV-x-ray lasers.

DOI: [10.1103/PhysRevA.00.003400](https://doi.org/10.1103/PhysRevA.00.003400)

Extreme ultraviolet (XUV) and x-ray coherent ultrashort light pulses can be generated in tabletop high-order harmonic generation (HHG) sources. Important applications suffer, however, from the low efficiency inherent to HHG processes. Different techniques have been investigated to overcome this low efficiency and important advances have been made [1–8]. Strong-field assisted XUV-x-ray parametric amplification processes in particular are investigated since they can produce avalanche effects in the medium that result in an exponential growth of the XUV-x-ray signal [9–15].

In previous research we addressed the influence of weak XUV radiation pulses of photon energies far from the ionization potential of the medium added to the strong IR driving pulse in HHG processes [11,13,14,16,17]. In particular, we showed that forward scattering can be largely enhanced when an XUV pulse is optimally synchronized with the IR pulse [11]. This theoretical prediction was soon corroborated by experiments [12] and a further detailed comparison between the theory and the experimental measurements showed good qualitative agreement [14]. Strong-field-mediated intrapulse x-ray parametric amplification (IXPA) processes were identified in [14] as decisive for the amplification at the single-atom level. The theory used in these previous studies was based on the strong-field approximation (SFA) [18]. It is therefore important to investigate the effect of the Coulomb potential of the atom, which is basically neglected in the SFA, on the amplification effects.

In this paper we perform time-dependent Schrödinger equation (TDSE) simulations in one dimension [19–21], which correctly consider the effect of the Coulomb potential, and demonstrate the IXPA effect using an *ab initio* theory by considering the Li^+ ion ($I_p = 75.6$ eV) as the amplifying medium. We have chosen to investigate optimal parameter ranges for IXPA in Li^+ for two reasons. First, parametric XUV-x-ray amplification effects were shown to benefit at the single-atom level by high gas ionization potentials [13]. Harmonics with large contrast were predicted compared to the broad spectrum obtained from the lower ionization potentials of gases as He. Second, a high ionization potential will also favor the avalanche effect produced by the parametric amplification processes [14], because those need a high

degree of synchronization between the XUV-x-ray and IR pulses, and this synchronization is optimally preserved by the lower plasma densities produced in higher ionization potential media. The results that we present might hence be a first valuable guide for future experiments. In our simulations, we use the soft-Coulomb potential

$$V(x) = -\frac{1}{\sqrt{x^2 + \beta^2}}, \quad (1)$$

where we use atomic units and $\beta^2 = 0.0252$. The IR field is a 800-nm wavelength, 4-fs \sin^2 pulse with a peak intensity of 2.5×10^{15} W/cm², which produces a maximum of $\approx 0.2\%$ ground-state depletion. The seed XUV-x-ray pulse is a \sin^2 pulse with a peak intensity of 2.5×10^{11} W/cm² and a duration of 200 as. IXPA processes are sensitive to the different parameters involved in the interaction process, such as the intensity of the IR pulse and the associated ground-state depletion, as well as the duration and the frequency content of the weak XUV-x-ray attosecond pulse. In what follows we study the influence of the XUV-x-ray central photon energy together with the delay with respect to the IR pulse on the IXPA effect at the single-atom level.

Figure 1 shows a basic scheme for IXPA processes. The low-frequency region of the XUV-x-ray weak pulse is absorbed by the electron, a first step that is assisted by the presence of the strong IR driving pulse that modifies the Coulomb potential of the ion; in a second step the electron is accelerated in the continuum by the IR electric field, reaching the necessary energy at the correct time so that a higher-frequency region of the XUV-x-ray pulse, optimally the central frequency region, stimulates the emission in a third step. The process is produced in the time scale of the duration of the XUV-x-ray pulse and benefits from the reshaping (chirp) during the interaction with the medium, as will be further detailed below.

In order to describe IXPA processes it is useful to compute the XUV-x-ray probe absorption spectrum, as described in [22], which represents the absorption of a weak probe pulse following a nonlinear excitation by one or more *pump* laser pulses

$$S(\omega) \propto \text{Im}[E_X^*(\omega)D_X(\omega)]. \quad (2)$$

To probe the state of our system due to the nonlinear excitation caused by the presence of both the strong IR field and the

*carles.serrat-jurado@upc.edu

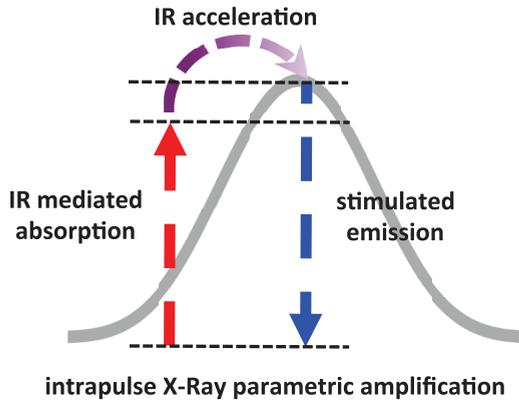


FIG. 1. Schematic of a strong-field-mediated IXPA.

93 weak XUV-x-ray pulse, we remove from the time-dependent
 94 dipole acceleration obtained in the TDSE integration the
 95 response of the system to the IR field alone and hence purely
 96 HHG processes are not included in the time-dependent dipole
 97 acceleration that plays the role of the pump in Eq. (2). We
 98 therefore define $D_X(\omega) = D(\omega) - D_{IR}(\omega)$, where $D(\omega)$ is the
 99 complete spectrum obtained from the excitation of the system
 100 and $D_{IR}(\omega)$ is the spectrum corresponding to the excitation
 101 of the system with the IR field alone. In this way, $S(\omega)$
 102 calculates the absorption that a weak probe pulse $E_X^*(\omega)$ would
 103 experience interacting with a system that has been pumped
 104 and presents the time-dependent dipole acceleration that gives
 105 the spectrum $D_X(\omega)$, which includes the linear response to
 106 the weak XUV-x-ray pulse. The probe pulse $E_X^*(\omega)$ in our
 107 calculations is the seed that we consider for the combined IR
 108 plus XUV-x-ray interaction.

109 Figure 2 shows the probe absorption spectra for the case of
 110 a 113-eV XUV pulse interacting with the ion together with the

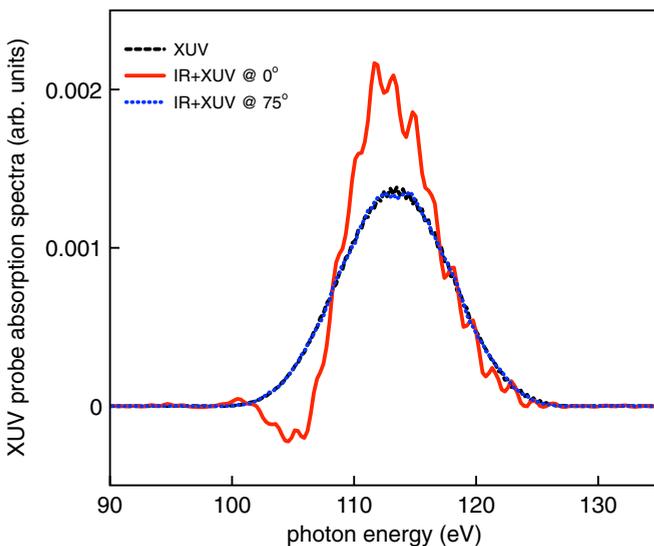


FIG. 2. XUV-x-ray probe absorption spectrum for a 113-eV XUV pulse interacting together with the IR pulse at a delay of 0° (red solid line), and 75° (blue dotted line), as indicated. The black dashed line shows the probe absorption spectrum in the case that the XUV pulse interacts alone with the ion.

111 IR pulse at two different delays. The red curve shows the case
 112 that the XUV pulse and the IR pulse are synchronized at a delay
 113 of 0° and the blue dotted curve shows the case for which the
 114 delay is equal to 75° . The results are compared to the case for
 115 which the weak XUV pulse interacts alone with the ion (black
 116 dashed curve), which gives a purely linear response. From the
 117 calculated probe absorption spectra (Fig. 2), it is clear that only
 118 certain delays between the IR and the XUV pulses can produce
 119 an effect different from the linear response. The red curve at
 120 a delay of 0° shows probe emission (negative values) at the
 121 lower frequencies of the XUV pulse, around 105 eV, which
 122 means that the probe pulse finds the system in an excited state,
 123 producing stimulated emission and hence probe gain at this
 124 frequency range. The red curve in Fig. 2 also shows probe
 125 absorption (positive values) around the central frequency of
 126 the XUV pulse that overcomes the probe absorption calculated
 127 for the purely linear case (black line), which means that the
 128 stimulated emission produced in this region by the IXPA
 129 process has depleted the excited state as compared to the linear
 130 case and therefore the probe absorption is higher. Indeed, the
 131 probe absorption spectral shape shown by the red line in Fig. 2
 132 is a signature of an intrapulse parametric process.

133 For a full characterization of the parametric processes we
 134 use two more tools. (i) We compare the spectra from the
 135 response of the system to the IR pulse alone, the XUV-
 136 x-ray pulse alone, and the combined IR plus XUV-x-ray
 137 pulses considering different delays between them. (ii) We
 138 perform a frequency-time analysis to determine at which times
 139 the signals are emitted. Furthermore, in the frequency-time
 140 analysis, we identify the IR-driven stimulated emission and
 141 absorption processes by removing the contributions from the
 142 signal produced by the IR pulse alone and the one produced
 143 by the XUV-x-ray pulse alone. As it will be further
 144 discussed below, this analysis allows us to determine the
 145 parameter regions where IR-driven parametric amplification
 146 in the spectral range of the seeded XUV-x-ray pulse occurs.

147 First, in order to determine the parameter ranges where
 148 nonlinear stimulated absorption occurs, we have calculated
 149 the integral of the negative regions of the XUV-x-ray probe
 150 absorption spectrum by varying both the XUV-x-ray central
 151 photon energy from 90 eV to 400 eV and the delay between the
 152 XUV-x-ray pulse and the IR field between -180° and 180° .
 153 The resulting map (Fig. 3) illustrates where the IXPA processes
 154 most probably arise, since nonlinear stimulated absorption is
 155 a necessary step to produce them. Note that Fig. 3 nicely
 156 reproduces the shape of the IR electric field as it was already
 157 observed in [14]. The particular cases shown in Fig. 2 hence
 158 correspond to the 113-eV photon energy value in Fig. 3 at the
 159 delays 0° and 75° .

160 Figure 4 shows the output spectra obtained for the particular
 161 case shown in Fig. 2 around 113 eV. The black dashed line
 162 in Fig. 4 corresponds to the spectrum obtained with only
 163 the strong IR pulse and the blue dotted line is the spectrum
 164 obtained with only the 113-eV XUV pulse. We note that the
 165 HHG yield produced by the Li^+ ion is about four orders
 166 of magnitude less than the linear response of the ion to the
 167 XUV pulse alone. The red line in Fig. 4 shows the spectrum
 168 obtained when both the XUV and IR pulses are sent together
 169 and synchronized at a delay of 0° . We can see that the spectral
 170 shape nicely follows the characteristics of an IXPA process,

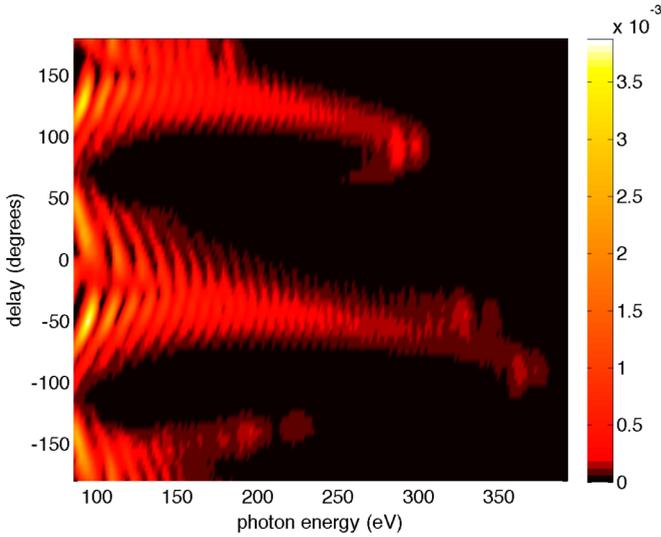


FIG. 3. Integrated negative regions of the XUV-x-ray probe absorption spectra as a function of the central photon energy of the XUV-x-ray pulse and the delay between the IR and XUV-x-ray pulse.

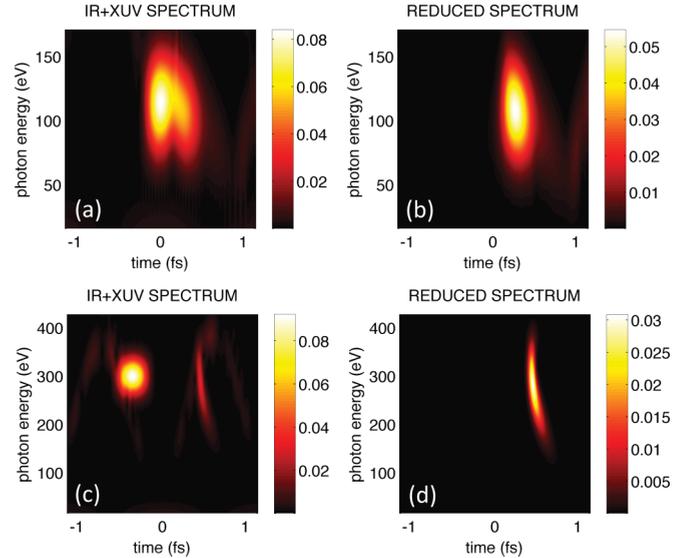


FIG. 5. Frequency-time analysis. (a) and (b) Spectra around 113 eV for a delay of 0° . (c) and (d) Spectra around 300 eV for a delay of -50° . The reduced spectra in (b) and (d) are obtained by subtracting the response of the IR pulse alone and the one of the XUV pulse alone from the total spectra, as explained in the text.

171 as it was also determined by the probe absorption spectrum in
 172 Fig. 2. We observe the absorption of the energy around 105
 173 eV and also the amplification of the signal beyond the linear
 174 response about the central frequency of the XUV pulse. When
 175 the XUV pulse is not optimally synchronized with the IR
 176 pulse, however, the nonlinear effects disappear and we obtain
 177 basically the same probe absorption spectrum (blue line in
 178 Fig. 2) and output spectrum (green line in Fig. 4) that are
 179 obtained by considering the interaction with the XUV pulse
 180 alone.

181 Figure 5 shows the frequency-time analysis of some of
 182 the spectra calculated by considering the results in Fig. 3.
 183 Figures 5(a) and 5(b) show the spectra at 113 eV for a
 184 delay of 0° and Figs. 5(c) and 5(d) correspond to the higher

photon energy of 300 eV and for a delay of -50° . Both cases 185
 correspond to areas in Fig. 3 where IXPA is expected to occur. 186
 Figure 5(a) shows the frequency-time spectra for the full IR 187
 plus XUV response around the 113-eV spectral region. In this 188
 figure we can observe both the effect of the linear response at 189
 the position where the XUV has been added, i.e., at a delay of 190
 0° ($t = 0$), and the nonlinear response around $t \approx 0.2$ fs. As 191
 already commented above, it is worth noting that the 200-as 192
 XUV pulse is chirped due to dispersion during the interaction 193
 with the ion and consequently time broadened, which allows 194
 the IXPA processes to be produced at $t \approx 0.2$ fs in this case. In 195
 Fig. 5(b) we plot the spectrum resulting from removing from 196
 the total response the one due to the IR pulse alone and the one 197
 due to the XUV pulse alone so that we remove the pure HHG 198
 contribution and the contribution of the linear response of the 199
 medium to the XUV pulse. We call it the reduced spectrum in 200
 Figs. 5(b) and 5(d) and hence it only shows the nonlinear 201
 contribution due to the combination of IR and XUV-x-ray 202
 assisted effects. We note that the amplified signal [Fig. 5(b)] 203
 is slightly down chirped. A similar behavior is shown in Figs. 5(c) 204
 and 5(d) by considering a higher-photon-energy x-ray seed 205
 pulse of 300-eV central photon energy. In this case the optimal 206
 delay, as calculated in Fig. 3, is about -50° ($t \approx -0.4$ fs). The 207
 nonlinear amplified IXPA signal is produced in this case at 208
 $t \approx 0.5$ fs, which again shows that the initial transform-limited 209
 x-ray seed is time broadened during the interaction with the 210
 Li^+ ion. 211

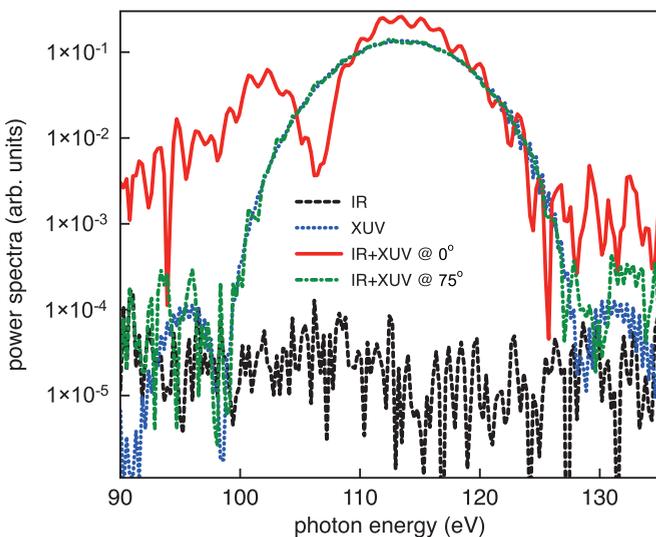


FIG. 4. Output spectra obtained from the Fourier transform of the time-dependent dipole acceleration, around the 113-eV region and for four different interaction cases, as indicated.

212 Finally, we address the effect of the added XUV-x-ray 212
 pulse (by taking as an example the 113 eV seed pulse) on 213
 the larger photon energy regions of the HHG spectrum. As 214
 shown in Fig. 6, a considerable increase of the HHG yield is 215
 obtained in the case of adding a seed pulse at 113 eV, with a 216
 plateau of harmonics that ends in this case at about 400 eV. 217
 We observe that this enhancement of the HHG yield depends 218

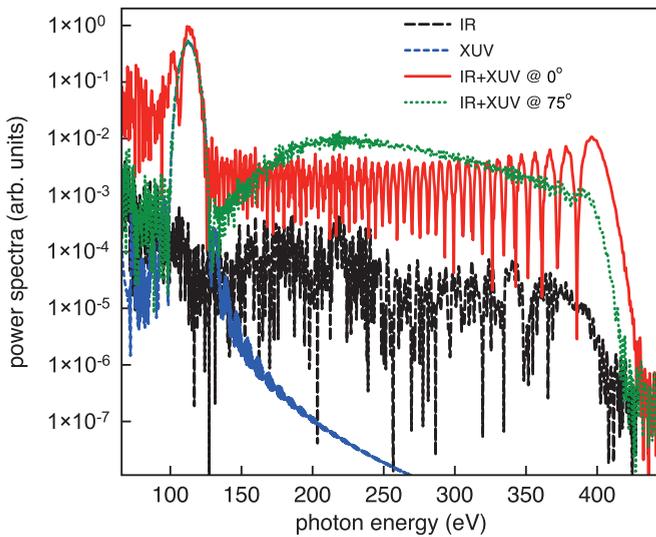


FIG. 6. Spectra obtained from the Fourier transform of the time-dependent dipole acceleration for the spectral region between 90 and 450 eV for four different interaction cases, as indicated.

the green line corresponds to a delay of 75° and shows a basically flat plateau. This behavior clearly corresponds to enhanced ionization phenomena that have been extensively studied [23–25]. Importantly, although it is not the central subject of this paper, we observe that the enhanced ionization effect is highly robust for the parameter ranges that have been here considered.

We have hence demonstrated IXPA processes in high-order harmonic generation at the single-atom level by using *ab initio* simulations and have determined optimal parameter ranges for IXPA processes to occur. Our study corroborates previous theoretical work performed under the SFA [11,13,14]. We conclude that several parameters are decisive and need to be balanced for IXPA processes to arise, such as the peak intensity of the driving IR pulse, the ionization potential of the medium, and the duration and spectral contents of the weak XUV–x-ray seed pulses. As already clear from previous studies, the delay between the weak seed and strong driving pulses is crucial for the amplification. The avalanche effects expected from IXPA processes when propagation and macroscopic effects are considered [14] result in exponential growth of the XUV–x-ray signal, which might be of importance for the implementation of intense tabletop XUV–x-ray lasers and also used as spectral narrow seeds for fully coherent plasma x-ray amplifiers.

Financial support from the Spanish Ministry of Economy and Competitiveness (FIS2014-51997-R) is acknowledged.

on the delay between the XUV and the IR pulse, as it is clear from Fig. 6. Indeed, the red line in Fig. 6 shows spectrally separated harmonics and corresponds to a delay of 0° , while

- [1] P. Agostini and L. F. DiMauro, *Rep. Prog. Phys.* **67**, 813 (2004).
- [2] T. Popmintchev, M.-C. Chen, D. Popmintchev, P. Arpin, S. Brown, S. Alisauskas, G. Andriukaitis, T. Balciunas, O. D. Mücke, A. Pugzlys, A. Baltuska, B. Shim, S. E. Schrauth, A. Gaeta, C. Hernández-García, L. Plaja, A. Becker, A. Jaron-Becker, M. M. Murnane, and H. C. Kapteyn, *Science* **336**, 1287 (2012).
- [3] C. Serrat and J. Biegert, *Phys. Rev. Lett.* **104**, 073901 (2010).
- [4] E. J. Takahashi, P. Lan, O. D. Mücke, Y. Nabekawa, and K. Midorikawa, *Nat. Commun.* **4**, 2691 (2013).
- [5] F. Brizuela, C. M. Heyl, P. Rudawski, D. Kroon, L. Rading, J. M. Dahlstrom, J. Mauritsson, P. Johnsson, C. L. Arnold, and A. L’Huillier, *Sci. Rep.* **3**, 1410 (2013).
- [6] S. Haessler, T. Balciunas, G. Fan, G. Andriukaitis, A. Pugzlys, A. Baltuska, T. Witting, R. Squibb, A. Zair, J. W. G. Tisch, J. P. Marangos, and L. E. Chipperfield, *Phys. Rev. X* **4**, 021028 (2014).
- [7] C. Jin, G. Wang, H. Wei, A.-T. Le, and C. D. Lin, *Nat. Commun.* **5**, 4003 (2014).
- [8] L. V. Dao, Kh. B. Dinh, and P. Hannaford, *Nat. Commun.* **6**, 7175 (2015).
- [9] J. Seres, E. Seres, D. Hochhaus, B. Ecker, D. Zimmer, V. Bagnoud, T. Kuehl, and C. Spielmann, *Nat. Phys.* **6**, 455 (2010).
- [10] J. Seres, E. Seres, and C. Spielmann, *Phys. Rev. A* **86**, 013822 (2012).
- [11] C. Serrat, *Phys. Rev. Lett.* **111**, 133902 (2013).
- [12] J. Seres, E. Seres, B. Landgraf, B. Ecker, B. Aurand, T. Kuehl, and C. Spielmann, *Sci. Rep.* **4**, 4234 (2014).
- [13] C. Serrat, D. Roca, and J. Seres, *Opt. Express* **23**, 4867 (2015).
- [14] C. Serrat, D. Roca, J. M. Budesca, J. Seres, E. Seres, B. Aurand, A. Hoffmann, S. Namba, T. Kuehl, and C. Spielmann, *Opt. Express* **24**, 8028 (2016).
- [15] T. Bredtmann, S. Chelkowski, A. D. Bandrauk, and M. Ivanov, *Phys. Rev. A* **93**, 021402(R) (2016).
- [16] C. Serrat, *Appl. Sci.* **2**, 816 (2012).
- [17] C. Serrat, *Phys. Rev. A* **87**, 013825 (2013).
- [18] M. Lewenstein, P. Balcou, M. Y. Ivanov, A. L’Huillier, and P. B. Corkum, *Phys. Rev. A* **49**, 2117 (1994).
- [19] G. N. Gibson and J. Biegert, *Phys. Rev. A* **78**, 033423 (2008).
- [20] E. Hijano, C. Serrat, G. N. Gibson, and J. Biegert, *Phys. Rev. A* **81**, 041401(R) (2010).
- [21] E. Hijano, C. Serrat, G. N. Gibson, C. Figueira de Morisso Faria, and J. Biegert, *J. Mod. Opt.* **58**, 1166 (2011).
- [22] S. Mukamel, in *Principles of Nonlinear Optical Spectroscopy* (Oxford University Press, Oxford, 1995), p. 103.
- [23] A. D. Bandrauk and N. H. Shon, *Phys. Rev. A* **66**, 031401(R) (2002).
- [24] J. Biegert, A. Heinrich, C. P. Hauri, W. Kornelis, P. Schlup, M. P. Ancombe, M. B. Gaarde, K. J. Schafer, and U. Keller, *J. Mod. Opt.* **53**, 87 (2006).
- [25] M. Tudorovskaya and M. Lein, *J. Mod. Opt.* **61**, 845 (2014).